

Anomaly detection in ZTF DR3

arXiv:2012.01419

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on the behalf of the SNAD+ team:

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DESC SN subgroup meeting, 2020.12.15



International research group for anomaly detection



2018, Moscow

<http://snad.space>

2020, Zoom



2019, Clermont-Ferrand

Methodology

Data

Zwicky Transient Facility Data Release 3:

- ~3 billion light curves in *gri* in 1020 overlapping fields, $\sim 2 \cdot 10^{11}$ detections
- If an object was observed in p passbands and f fields it will be represented as $p \cdot f$ individual light curves with different object identifiers (OIDs)

Our data set:

- The first 9 months of the survey due to better cadence
- 3 fields and *zr* passband only, ~2.25 million light curves total

Field name	ZTF field	Centre (α , δ)	Centre (l , b)	FOV	N	Object count
M 31	Field 695, ccdid 11	10° 72, 41° 34	121° 20, -21° 51	3 deg ²	≥ 100	57 546
DEEP	Field 795	246° 72, 54° 88	83° 91, 42° 45	55 deg ²	≥ 100	406 611
DISK	Field 807	19° 99, 62° 05	126° 27, -0° 63	55 deg ²	≥ 500	1 790 565

Pre-processing

Feature extraction + normalisation

- 42 light-curve features, e. g. Lomb-Scargle Periogram-based features, magnitude statistics, irregular time-series features like η^e or Stetson K
- New feature extraction code written on Rust:
<https://github.com/hombit/light-curve>

Has Python pip-installable binding `light-curve-python`, which is an order magnitude faster than popular Python codes like Feets.

Currently it is a part of the ANTARES broker pipeline for each new ZTF alert

- Standard scaling: removing the mean and scaling to unit variance

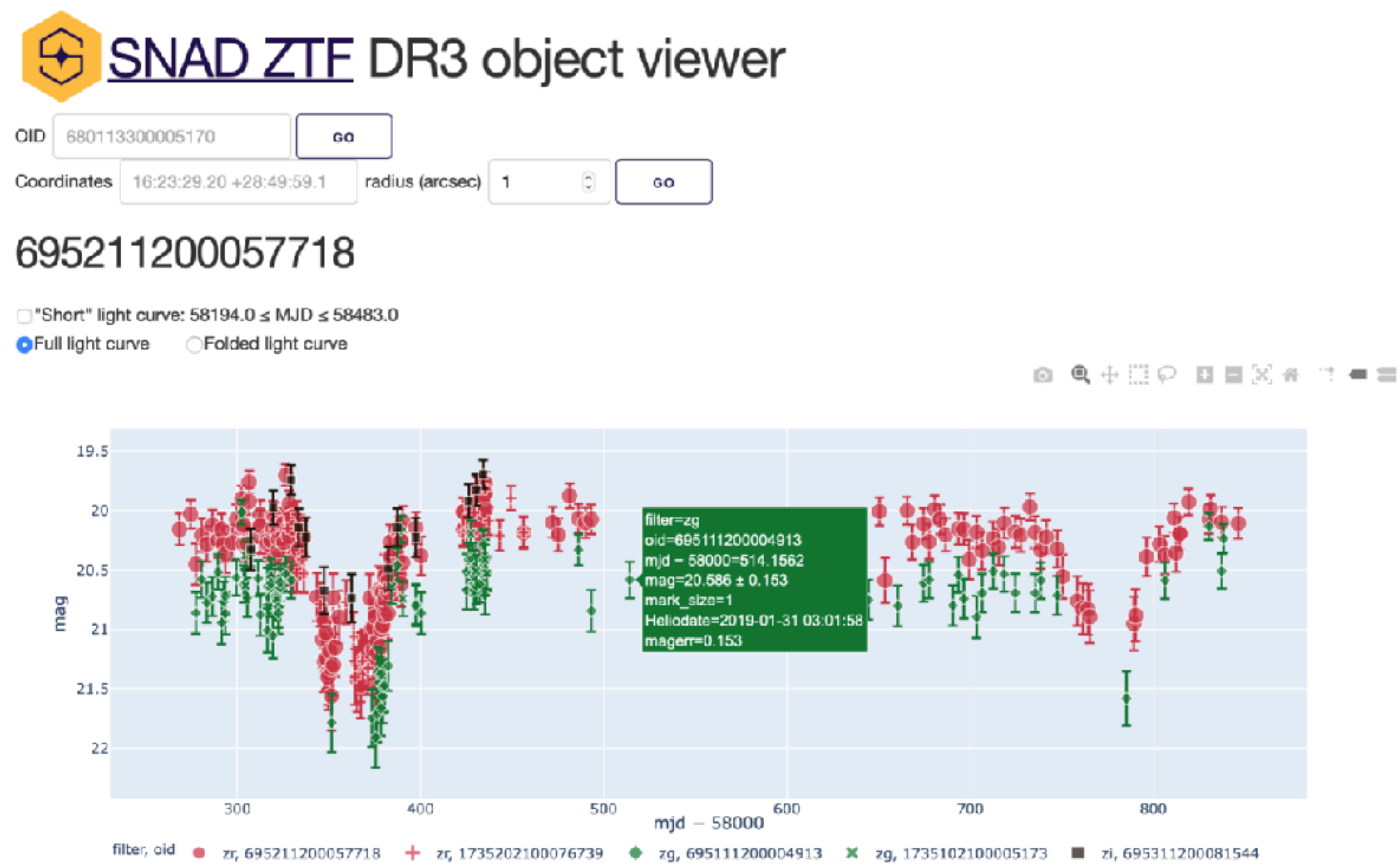
Anomaly Detection Pipeline

<https://github.com/snad-space/zwad>

- Def. *Outlier* is an object located in a sparse region of the feature space
- Def. *Anomaly* is an astrophysical source having unusual properties for its class or a representative of some rare class
- Four outlier detection algorithms: Isolation Forest, Gaussian Mixture Model, Local Outlier Factor, One-class Support Vector Machine
- Each field data are a separate data set
- Each algorithms is asked for 40 objects with the greatest outlier score for each field.
- Total 277 objects are examined using catalogs, literature, simulations, photometry redoing, and additional photometric and spectroscopic observations

SNAD ZTF web-viewer

<https://ztf.snad.space>



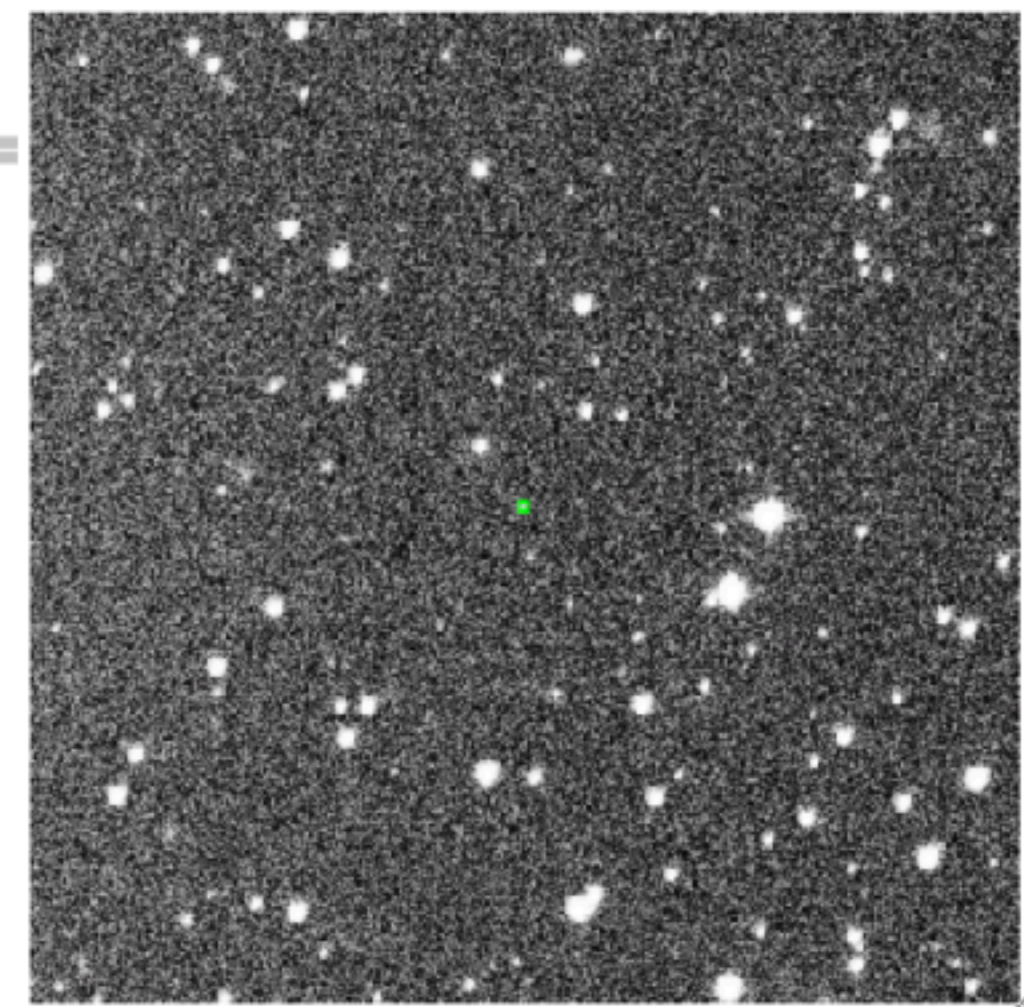
Download [PNG](#), [PDF](#)

Summary

Name: [DBL2003] 12 (27.159" [Simbad](#)), 381257684160228608 (0.168" [Gaia DR2 Distances](#))

Type: C* (27.159" [Simbad](#))

Distance: 1150.29 pc (0.168" [Gaia DR2 Distances](#))

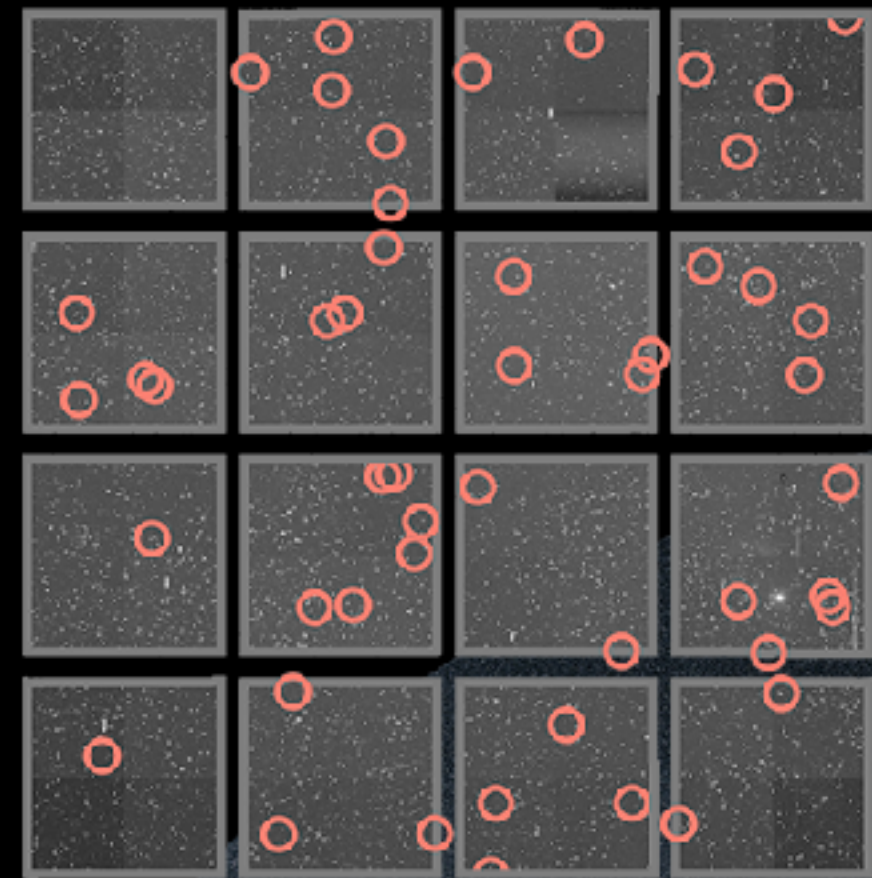


[Download FITS](#)

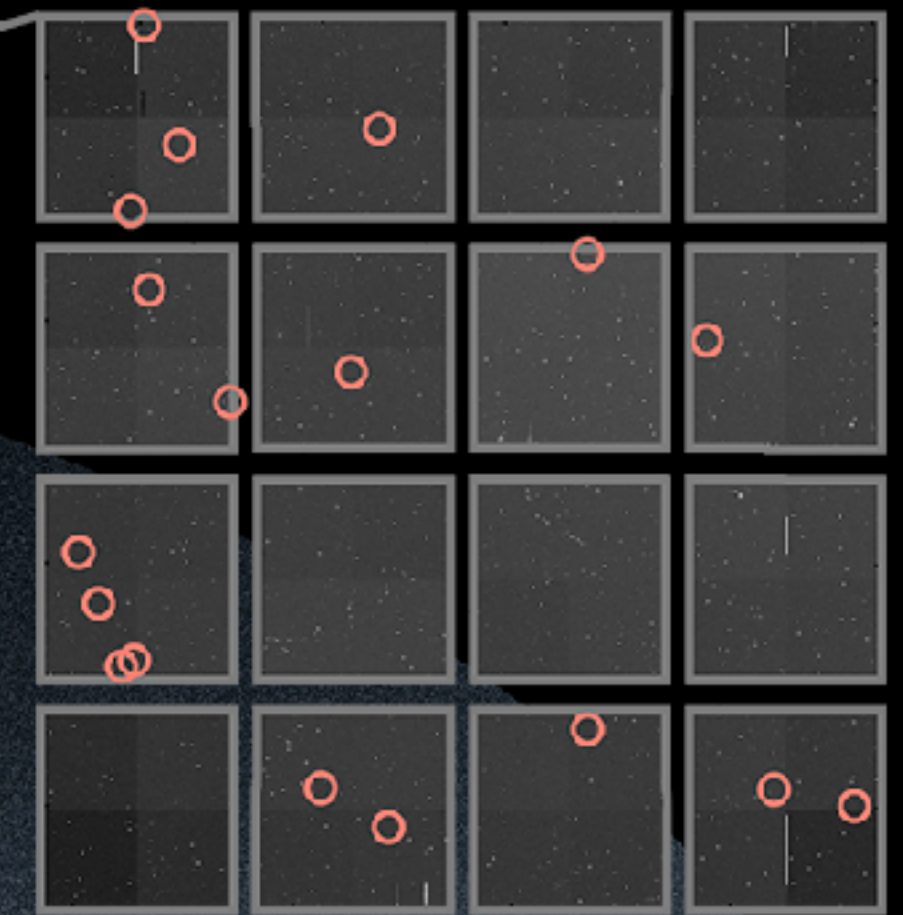
Aladin



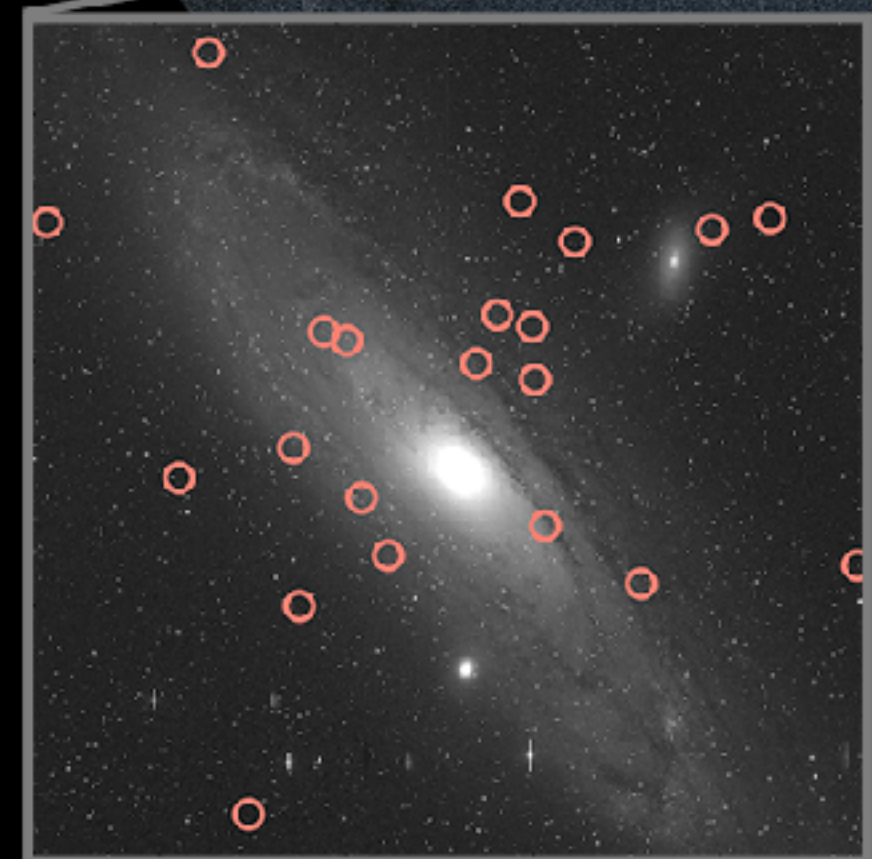
Anomaly Candidates — 89 / 277



DISK



DEEP



M31



**ANOMALY
CANDIDATE**

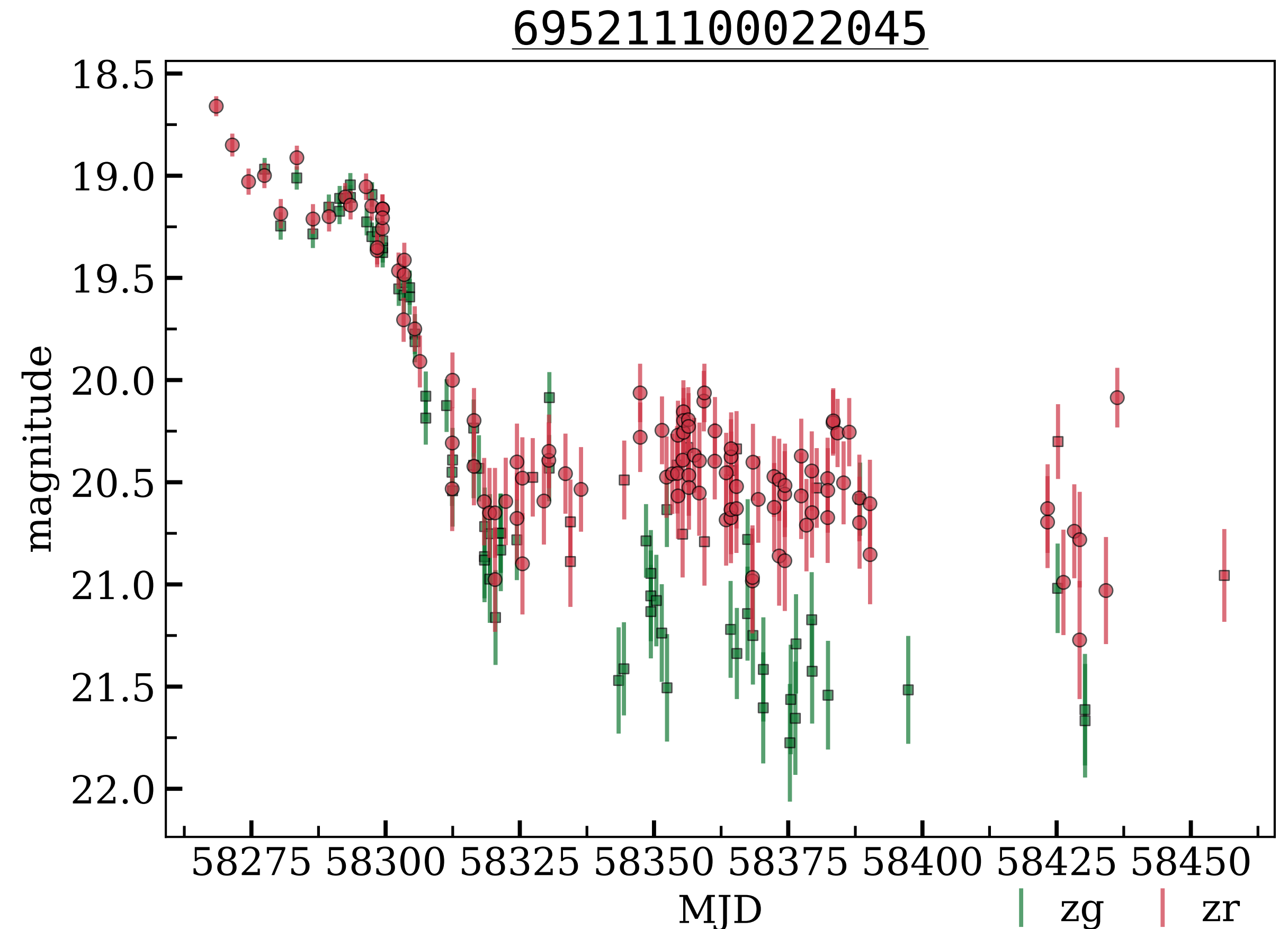


SNAD

Recurrent Nova Candidate

AT 2017ixs from M31 field, it had had two outbursts in 2017, before ZTF

- Red circles are the considering by the anomaly detection pipeline light curve
- Squares are other light curves of the same source but in different fields and/or passbands

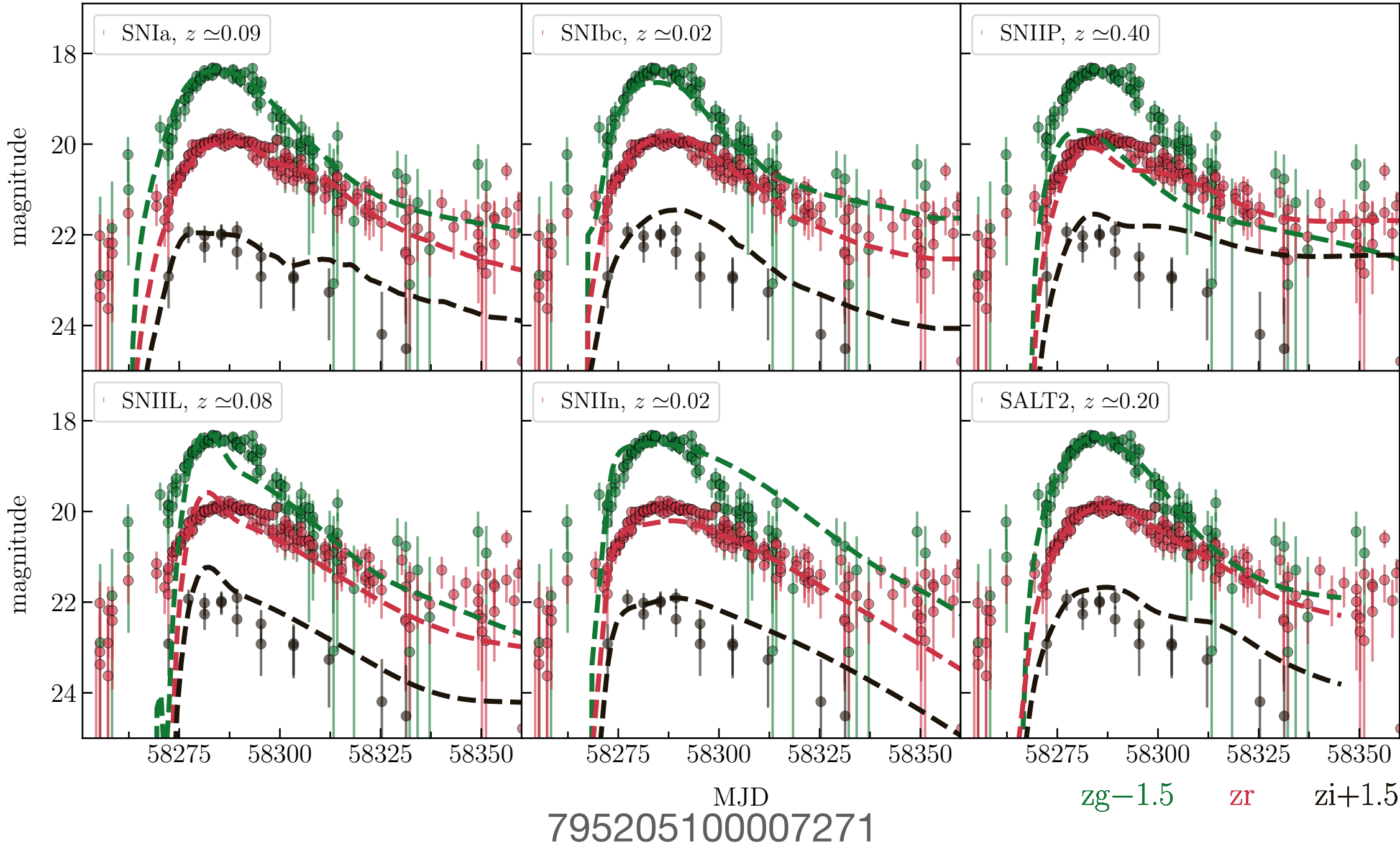
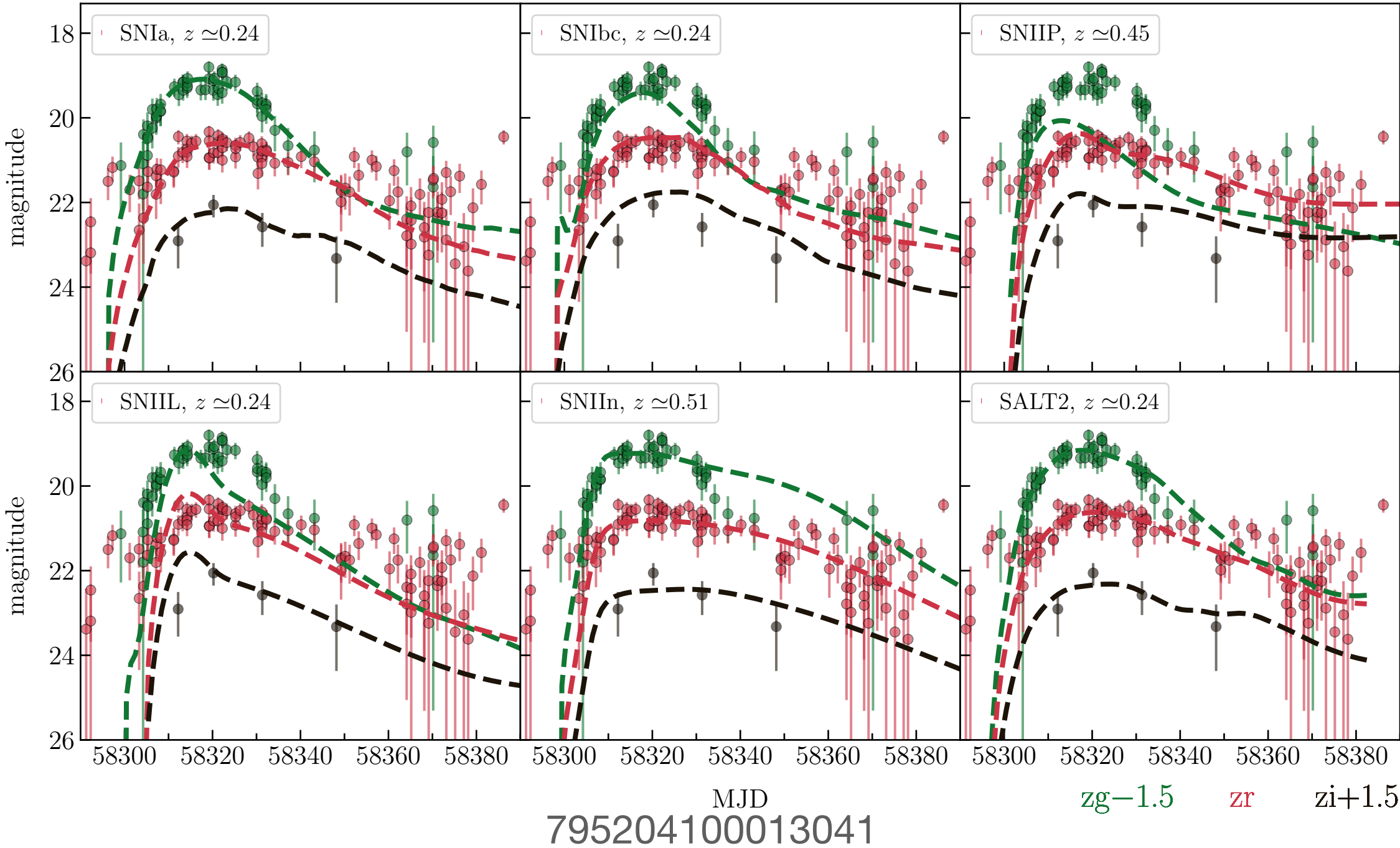


Supernova Candidates

Six candidates from DEEP field (400 000 objects), only one is in TNS

Table 2. Results of the light curve fit with the SALT2 model for supernova candidates from the DEEP field.

OID	Host galaxy*	z_{ph}	z	t_0	x_1	c	Comments [†]
795202100005941/ZTF18aanbnjh	SDSS J163437.92+521642.2	0.424 ± 0.103	—	—	—	—	Blazar
795204100013041/ZTF18abgvctp	SDSS J160913.83+521251.3	0.375 ± 0.138	~ 0.24	58320.9336 ± 0.4389	1.71 ± 0.51	-0.044 ± 0.035	—
795205100007271/ZTF18aayatjf	—	—	~ 0.20	58285.8334 ± 0.1810	-0.54 ± 0.18	-0.075 ± 0.021	SN Ia
795209200003484/ZTF18abbpebf	—	—	~ 0.11	58299.7269 ± 0.0008	0.60 ± 0.12	-0.013 ± 0.012	SN Ia
795212100007964/ZTF18aanbksg	SDSS J161144.90+555740.7	0.288 ± 0.122	~ 0.18	58214.4470 ± 0.0002	0.40 ± 0.20	-0.282 ± 0.020	Blazar
795213200000671/ZTF18aaincjb	—	—	—	—	—	—	AGN-I

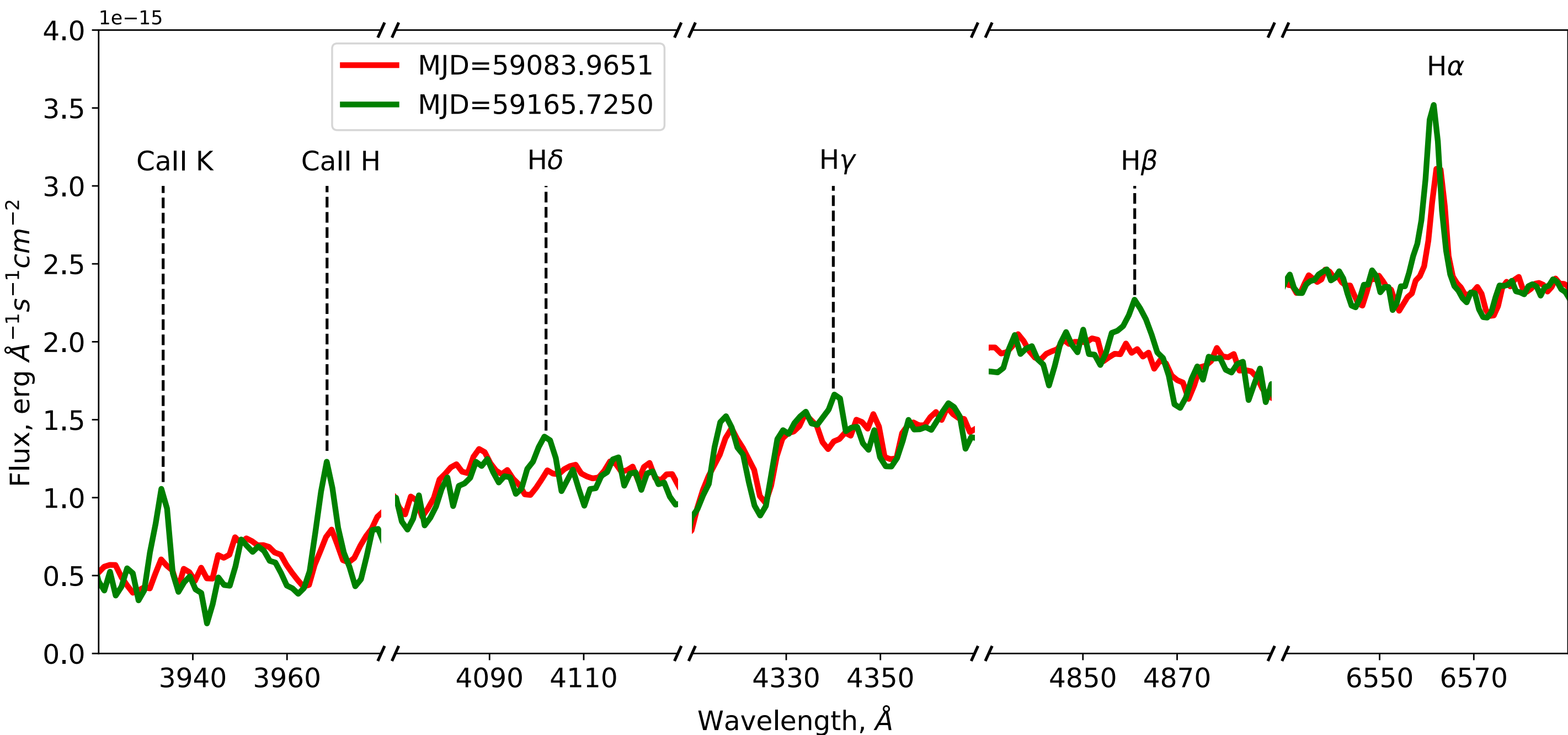
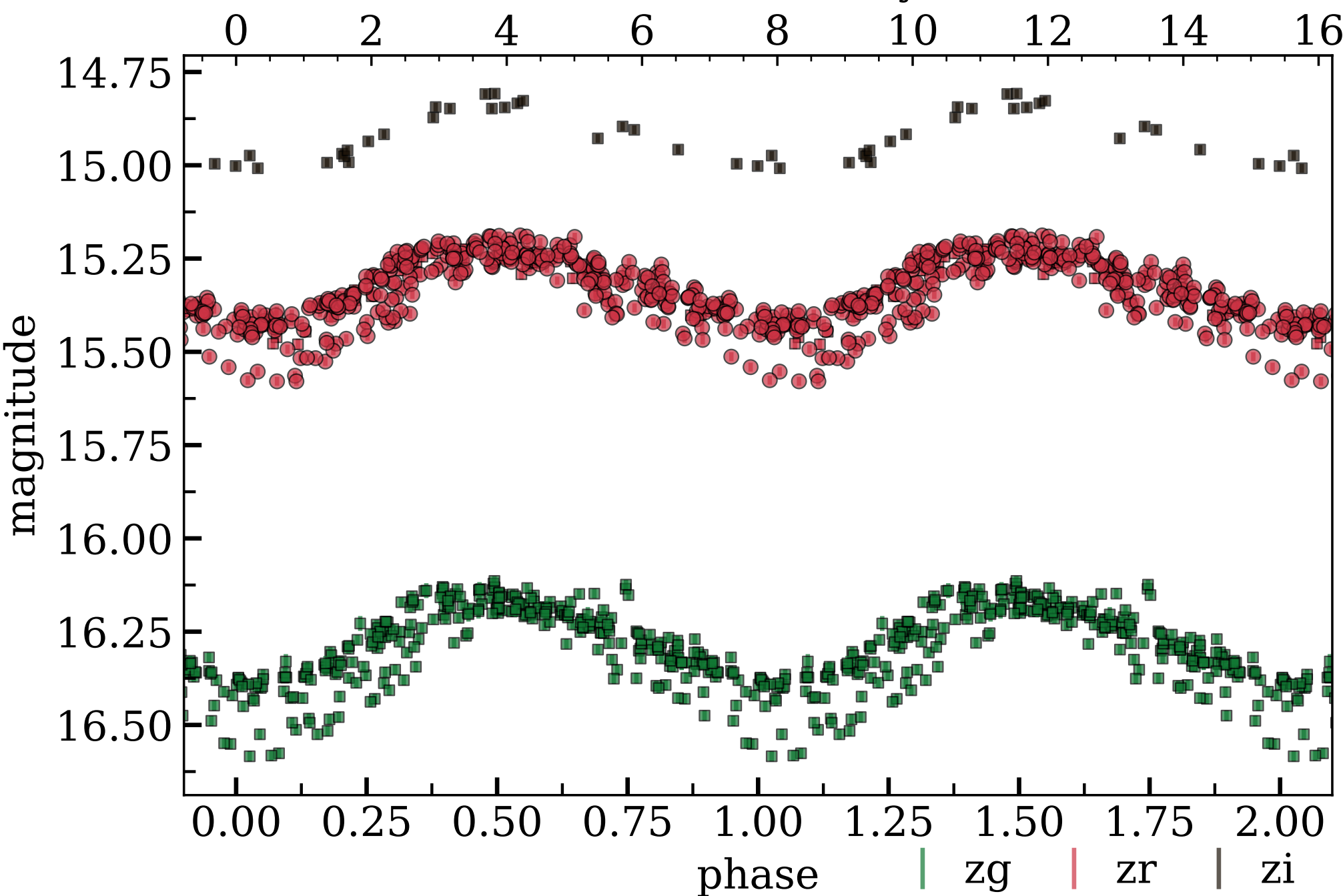


Classification of RS CVn (binary w/ spots)

Our spectra + period change + flare activity

695211200019653, $P = 7.715$ days

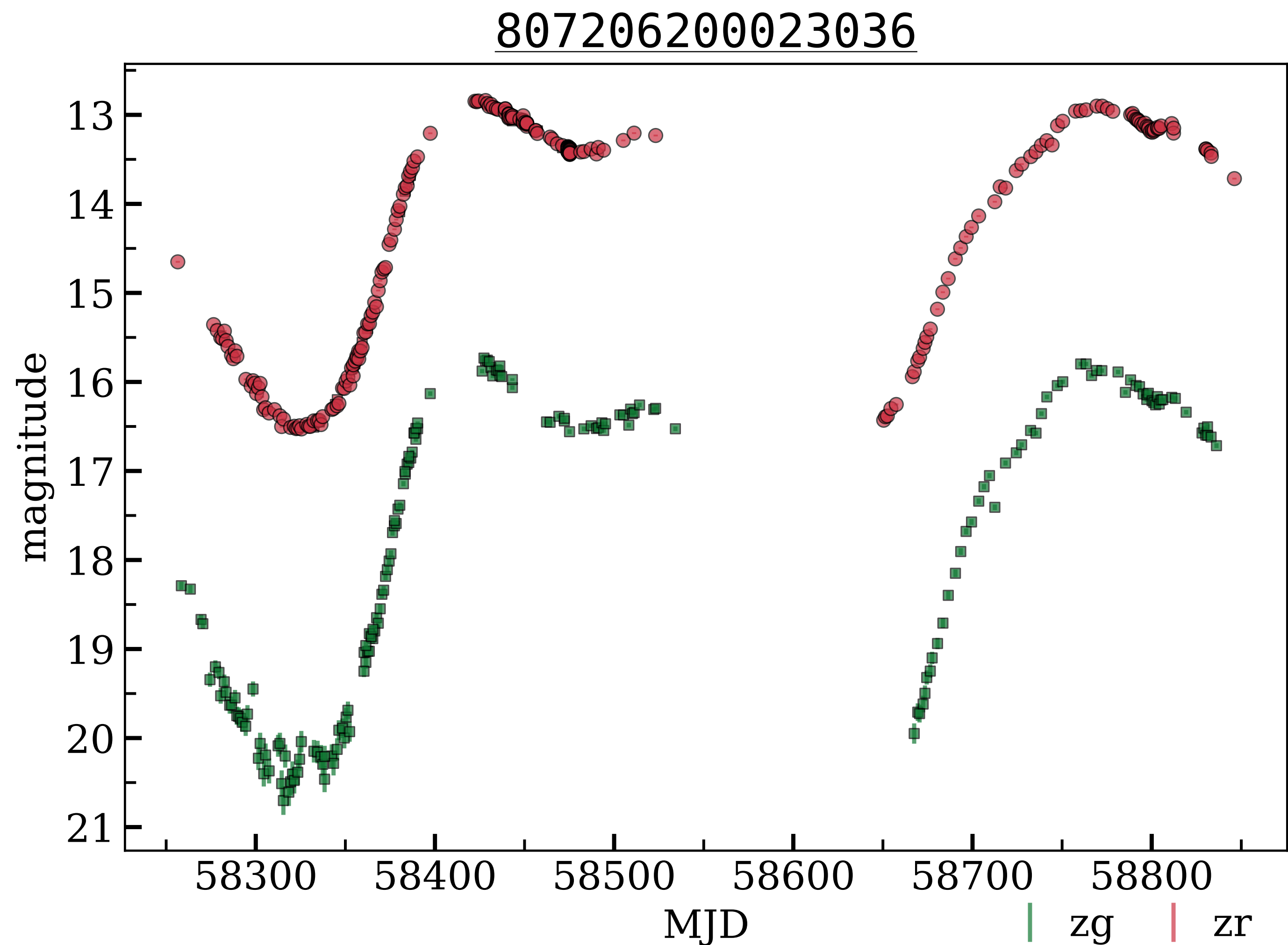
Folded time, days



2.5m CMO SAI MSU spectra of the object
at different phases of the orbital cycle

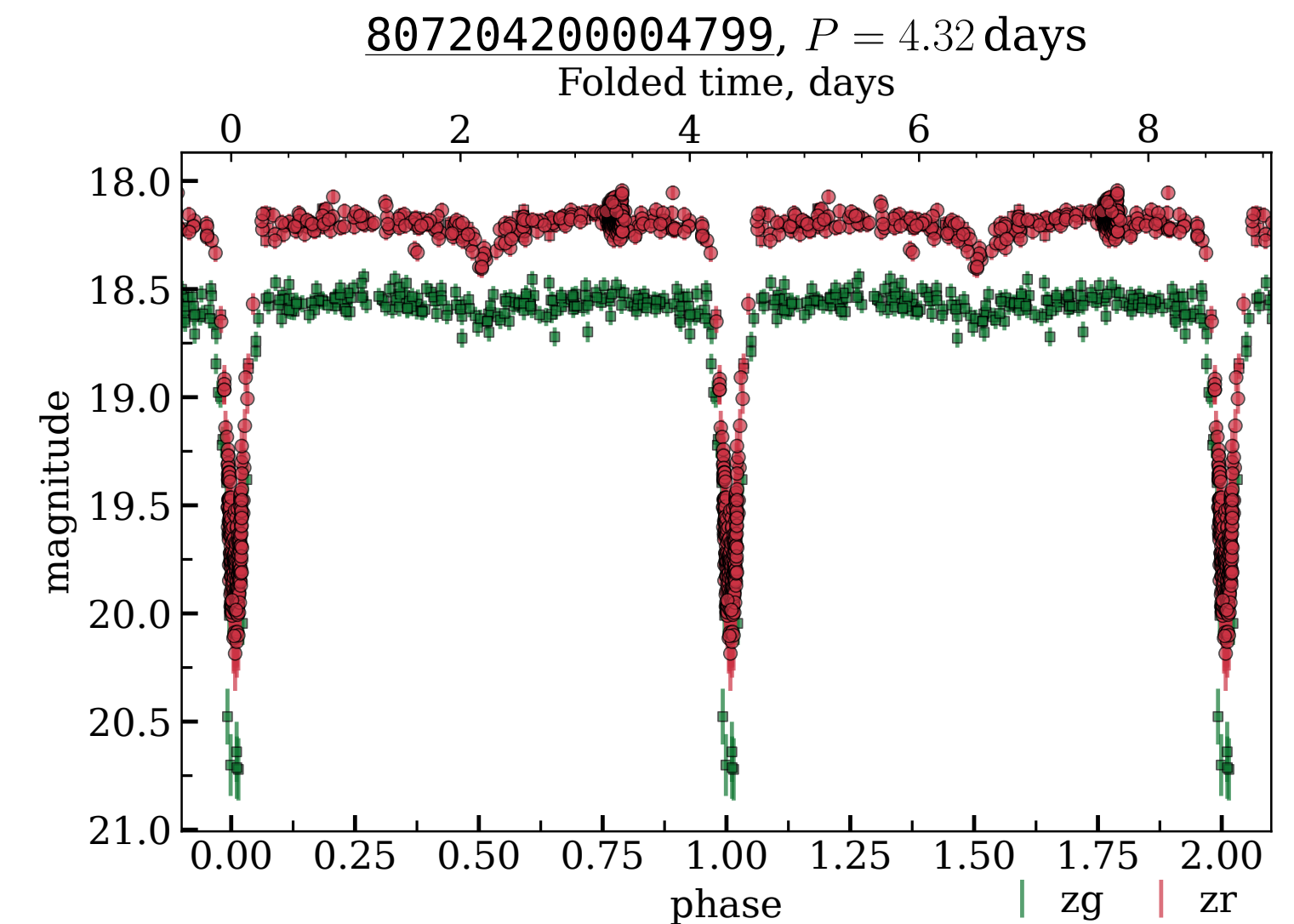
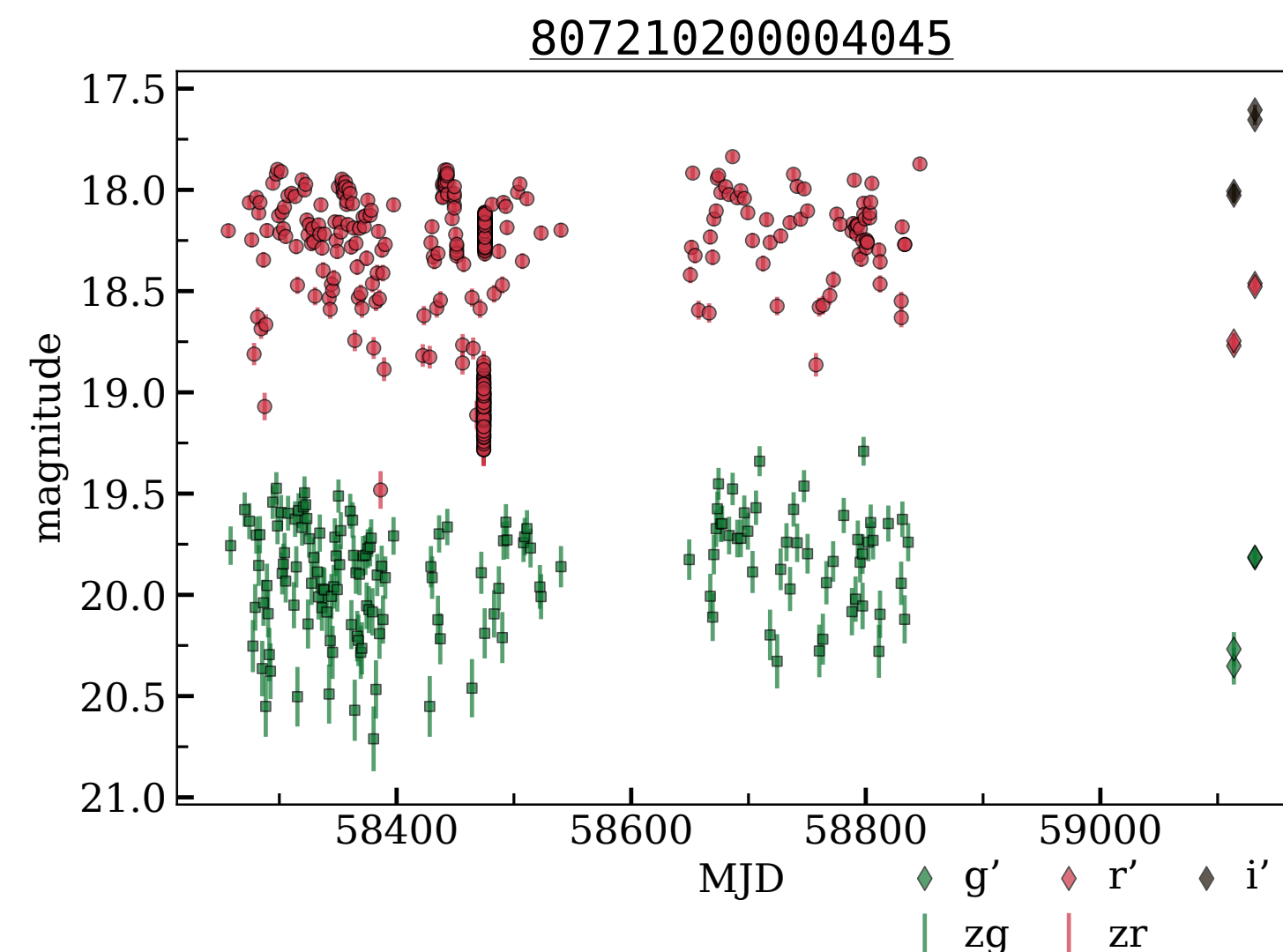
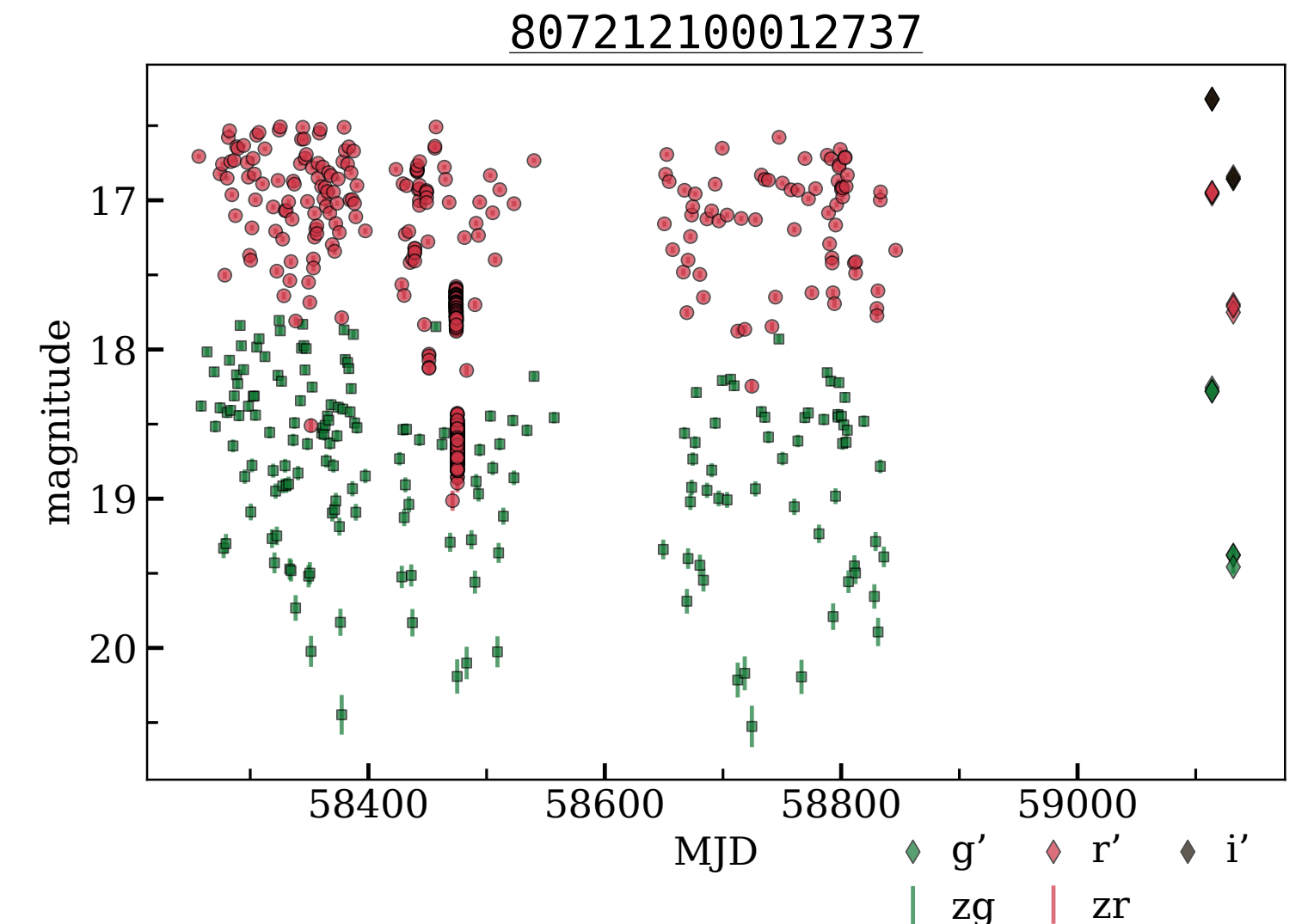
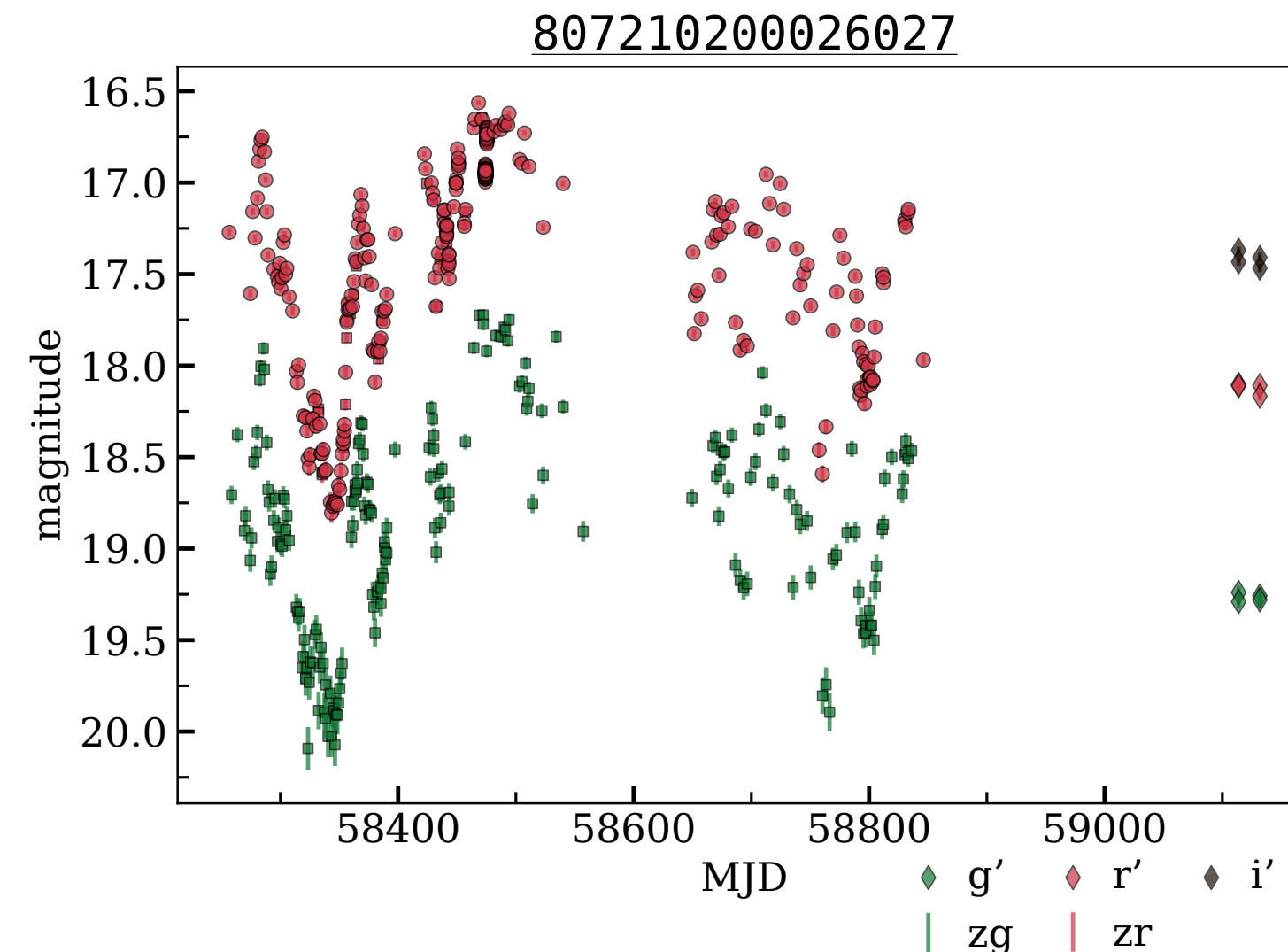
Mira Binary Candidate

Light curve may indicate the presence of a companion



Non-catalogued Sources — 23 / 277

- Diamonds are our observations (the right group of observations)
- Red circles are the considering by the anomaly detection pipeline light curve
- Squares are other light curves of the same source



Bogus Light Curves — 188 / 277

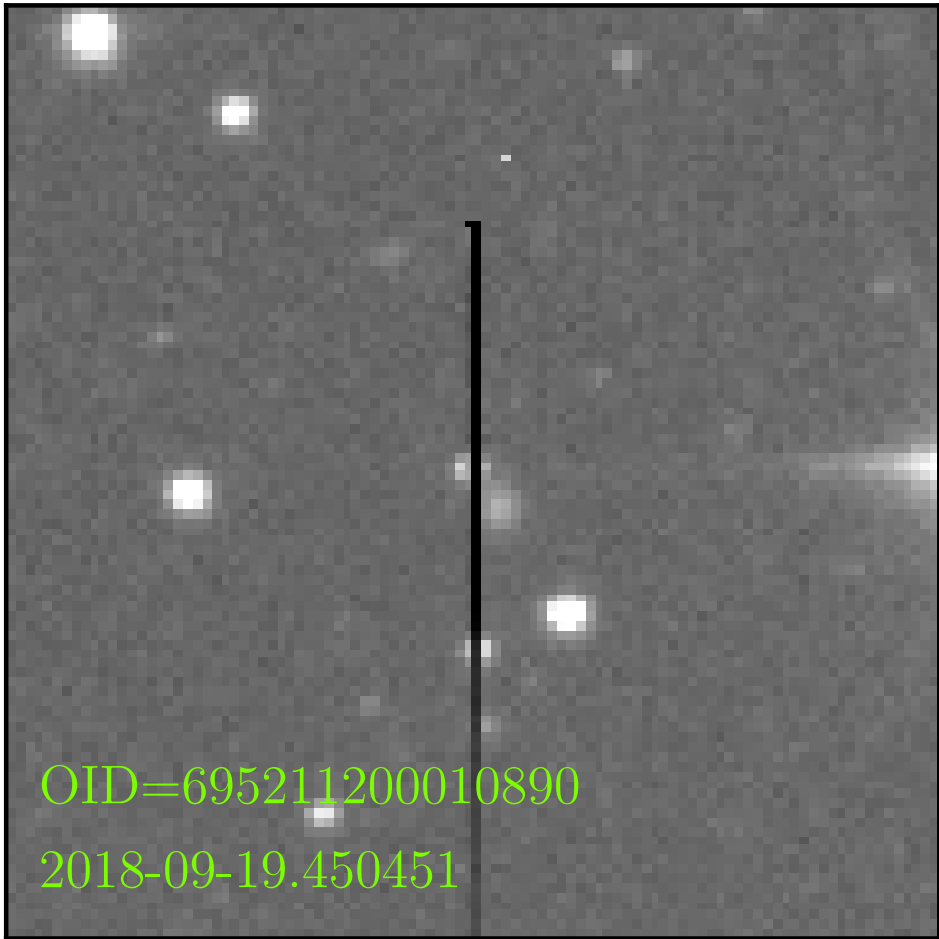
Bogus Detections

The objects are in the frame centers

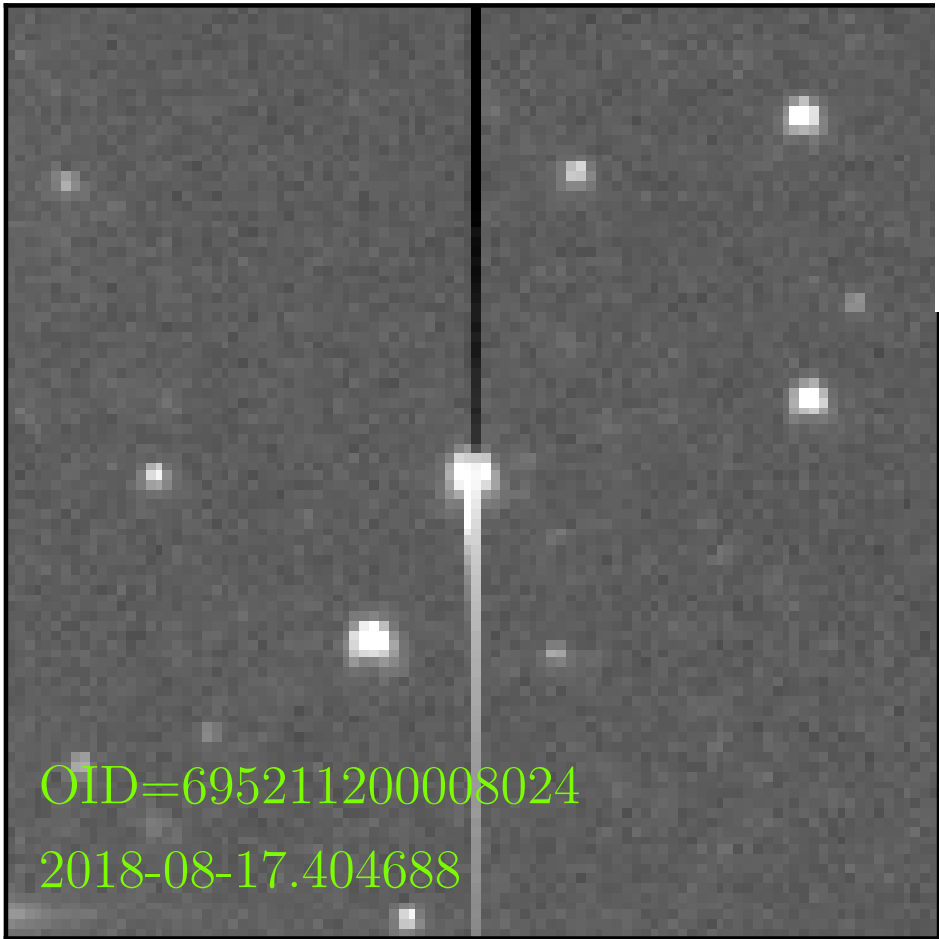
Satellite or Plane Track



Bad Column



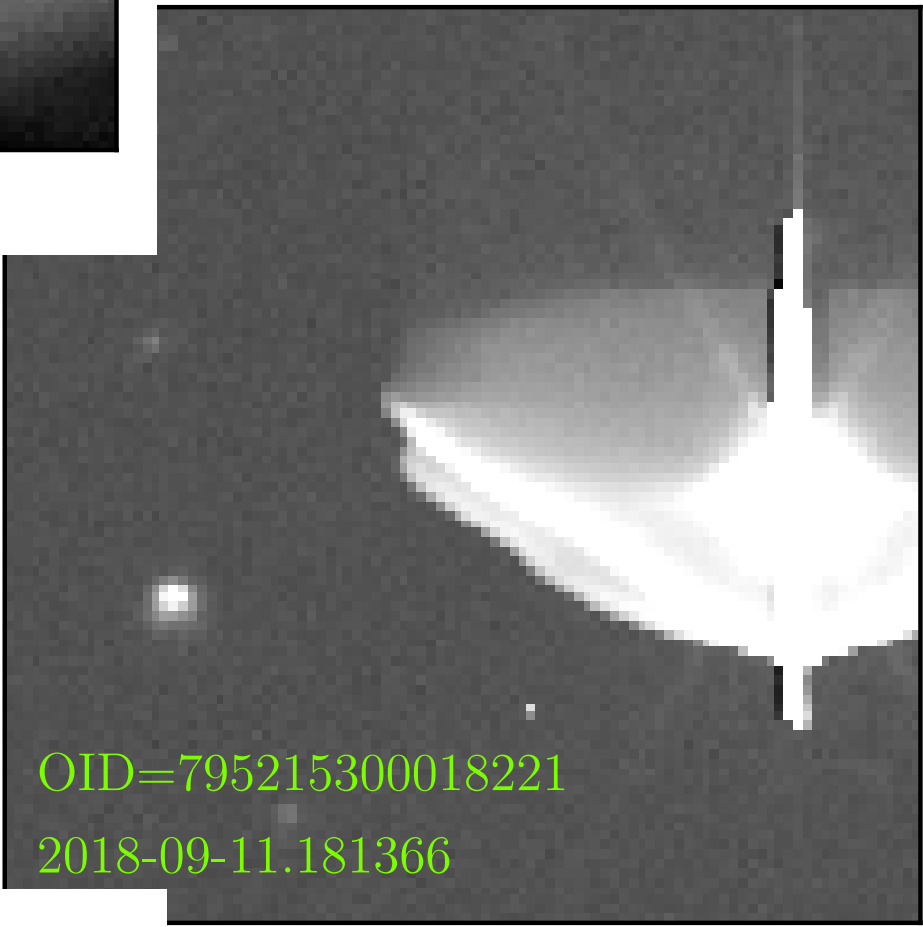
Bad Column



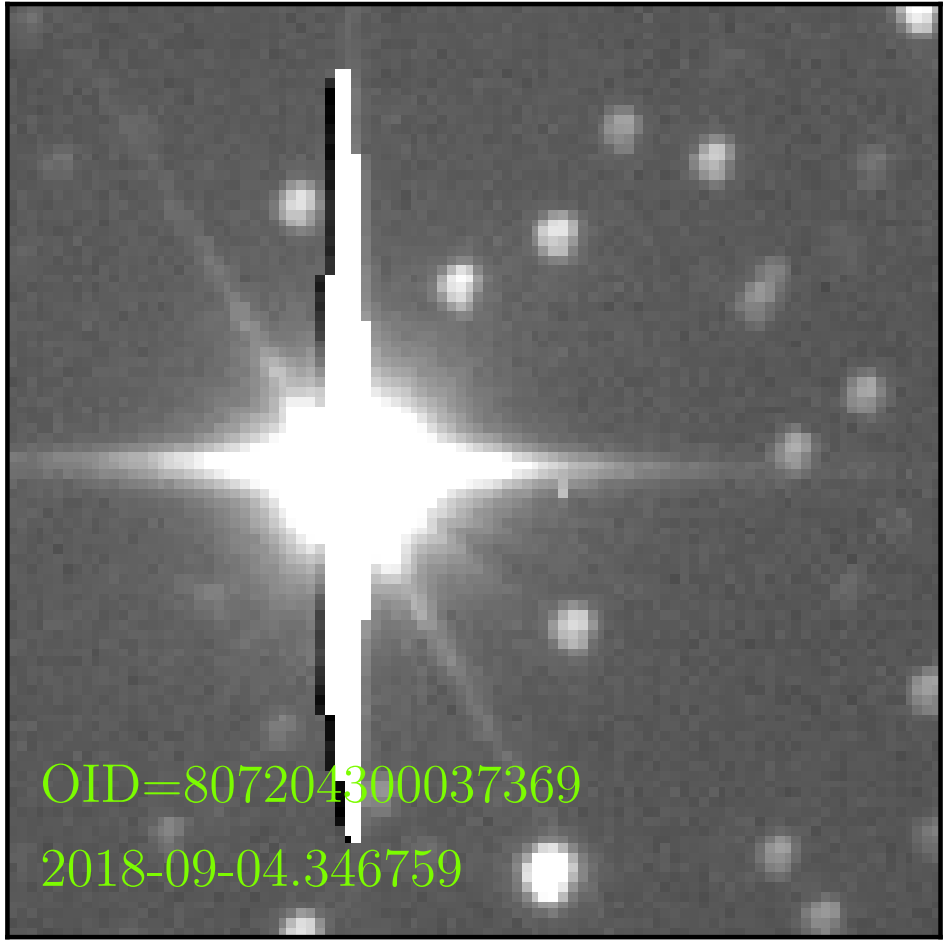
Close to M31 Centre



Ghost



Diffraction Spike



Cosmic Ray



Defocusing

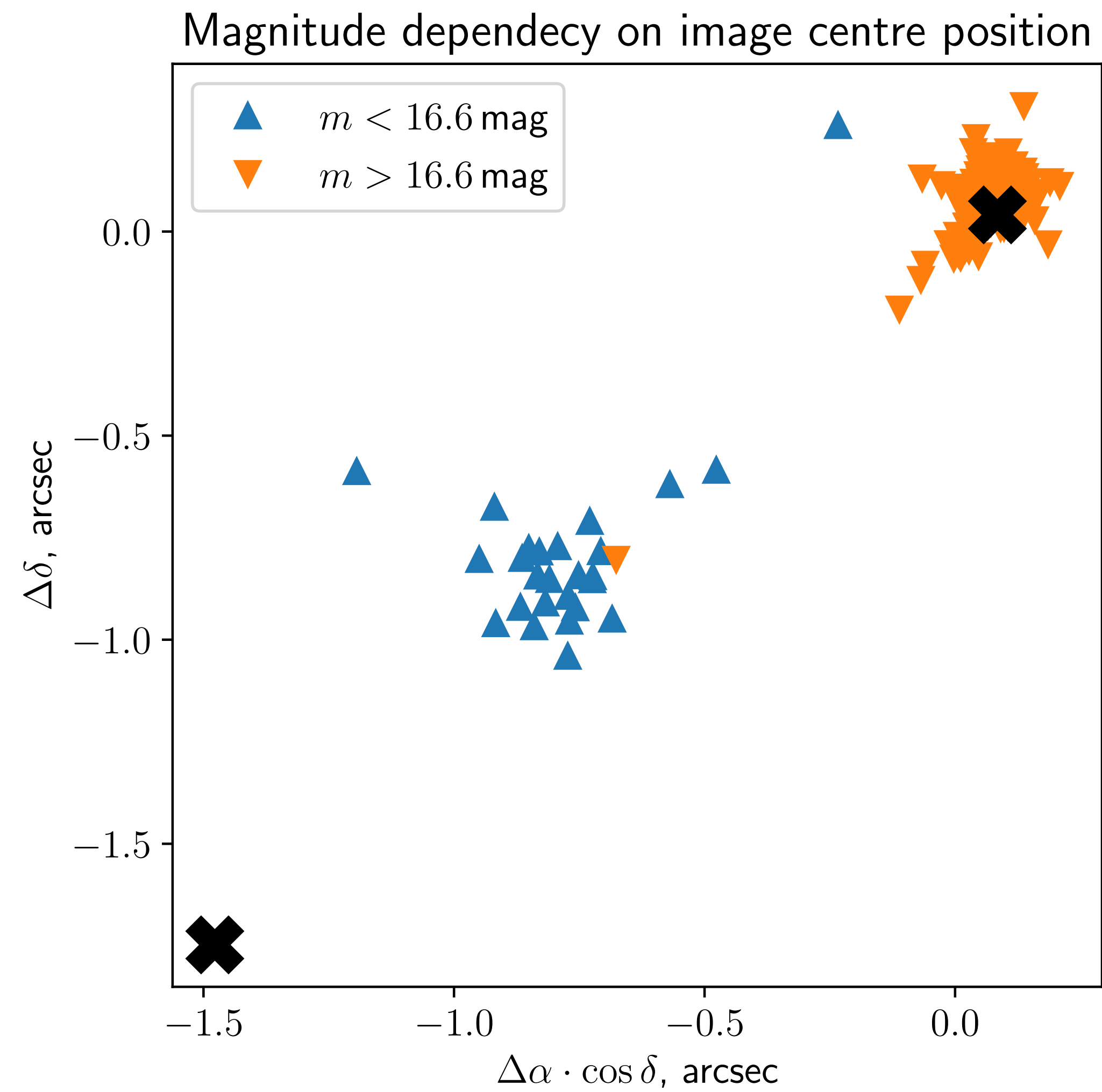
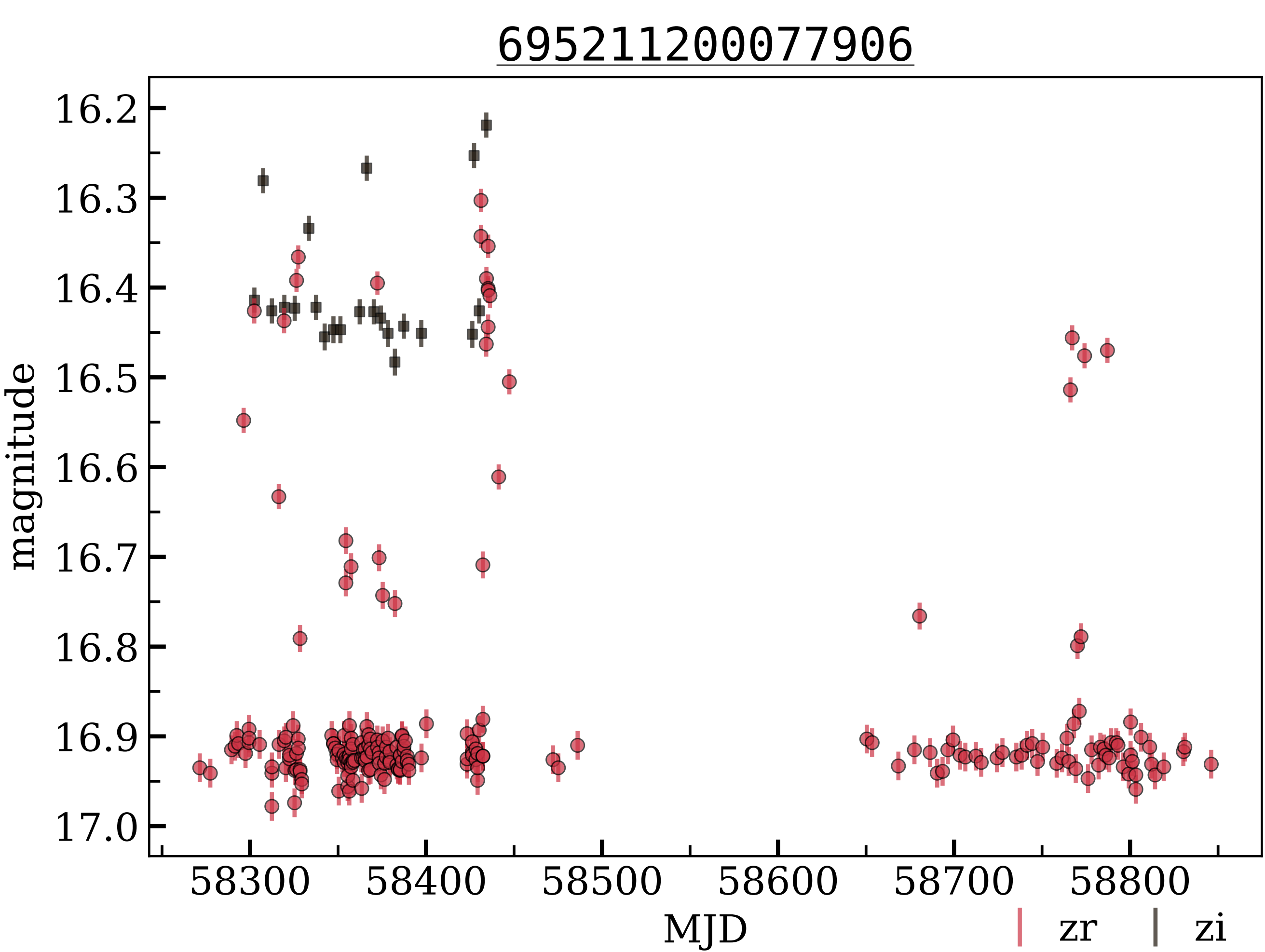


Ghost



Double Stars Defocusing

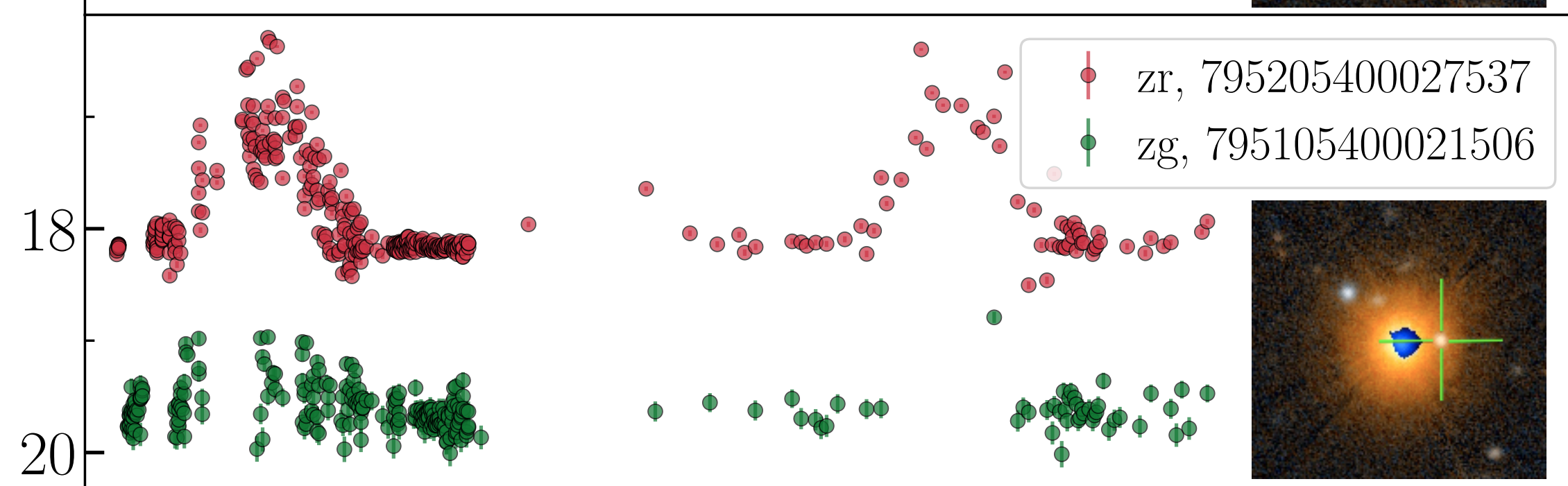
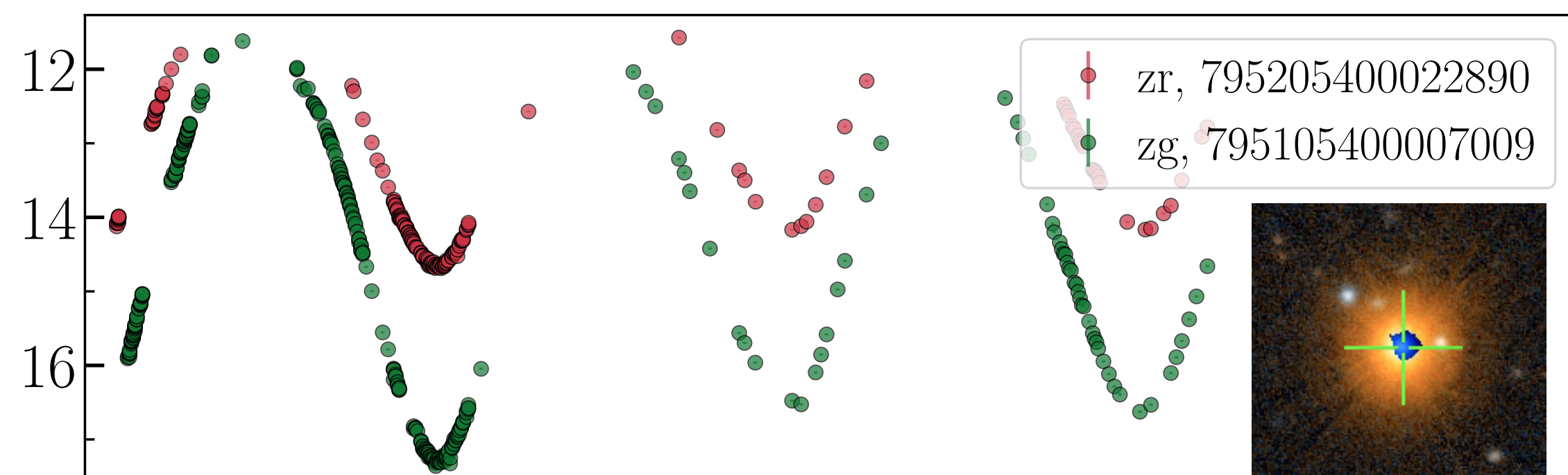
If a separation is about 2" (typical FWHM) then defocusing can cause the false variability



IW Dra "Echos"

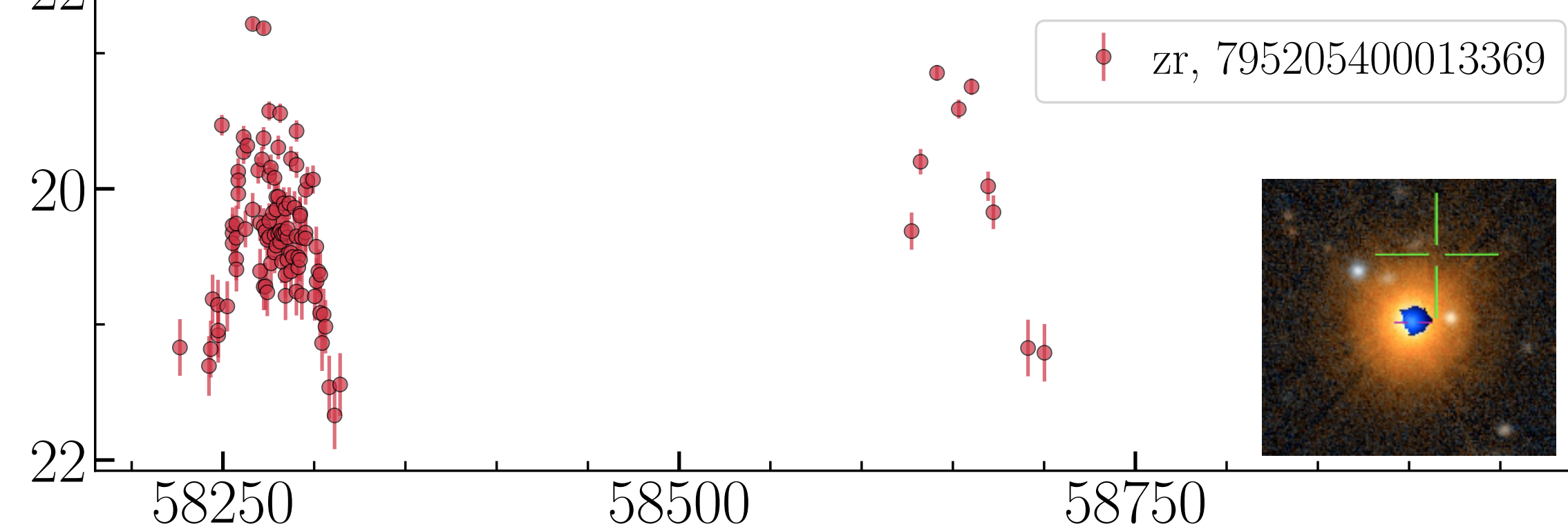
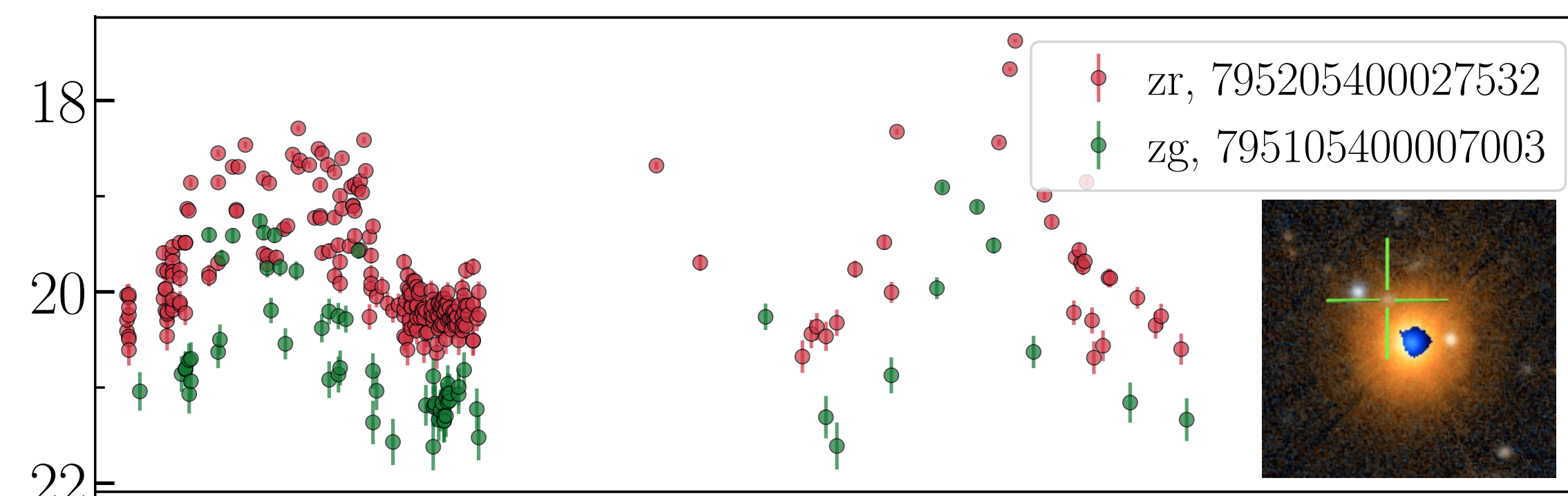
All four objects are found by the outlier detection pipeline

Mira-like IW Dra



Echo 1

Echo 2



Echo 3

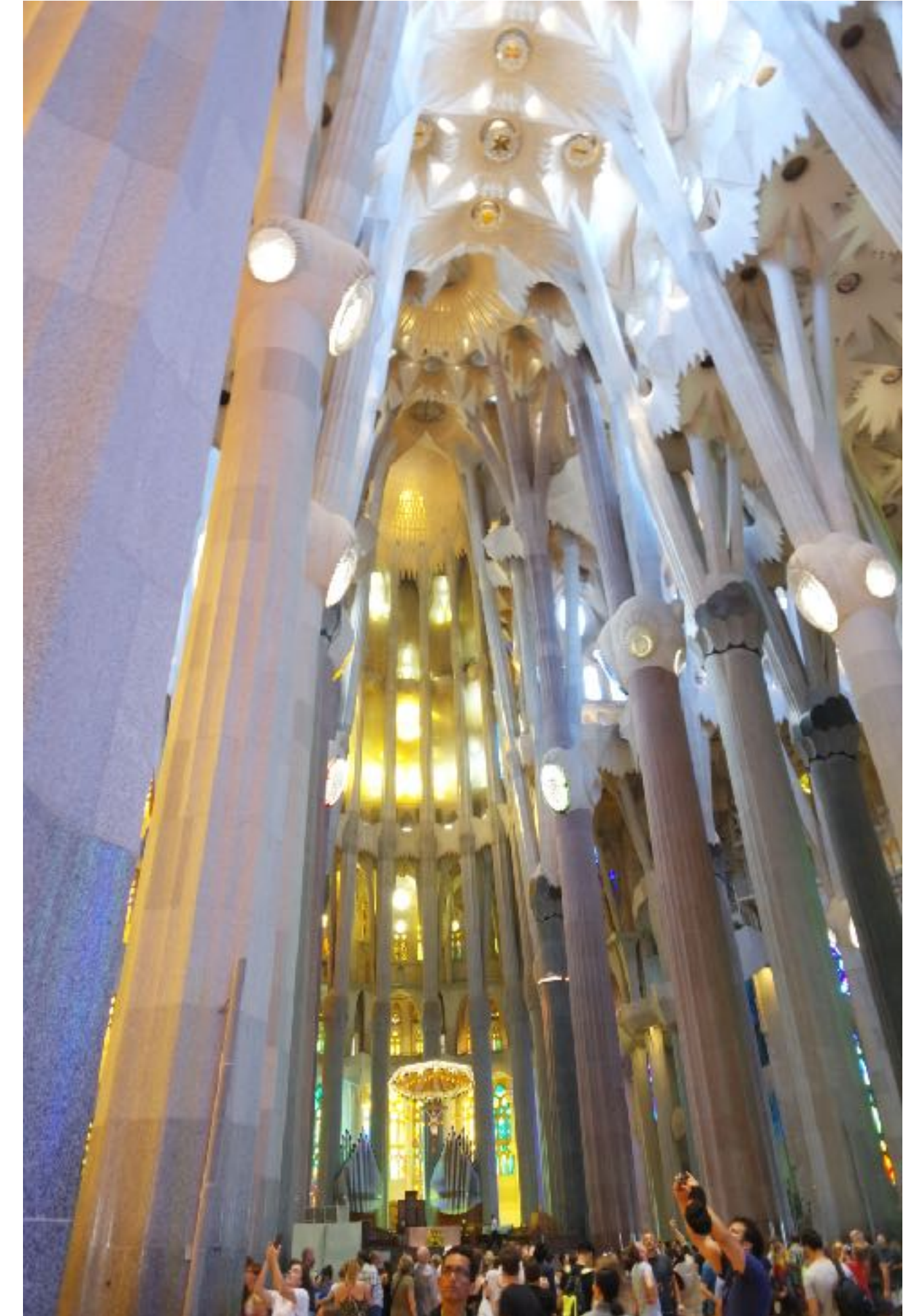
MJD

Barcelona

I bet you've never been to Barcelona



m0n5t3r @ Flickr

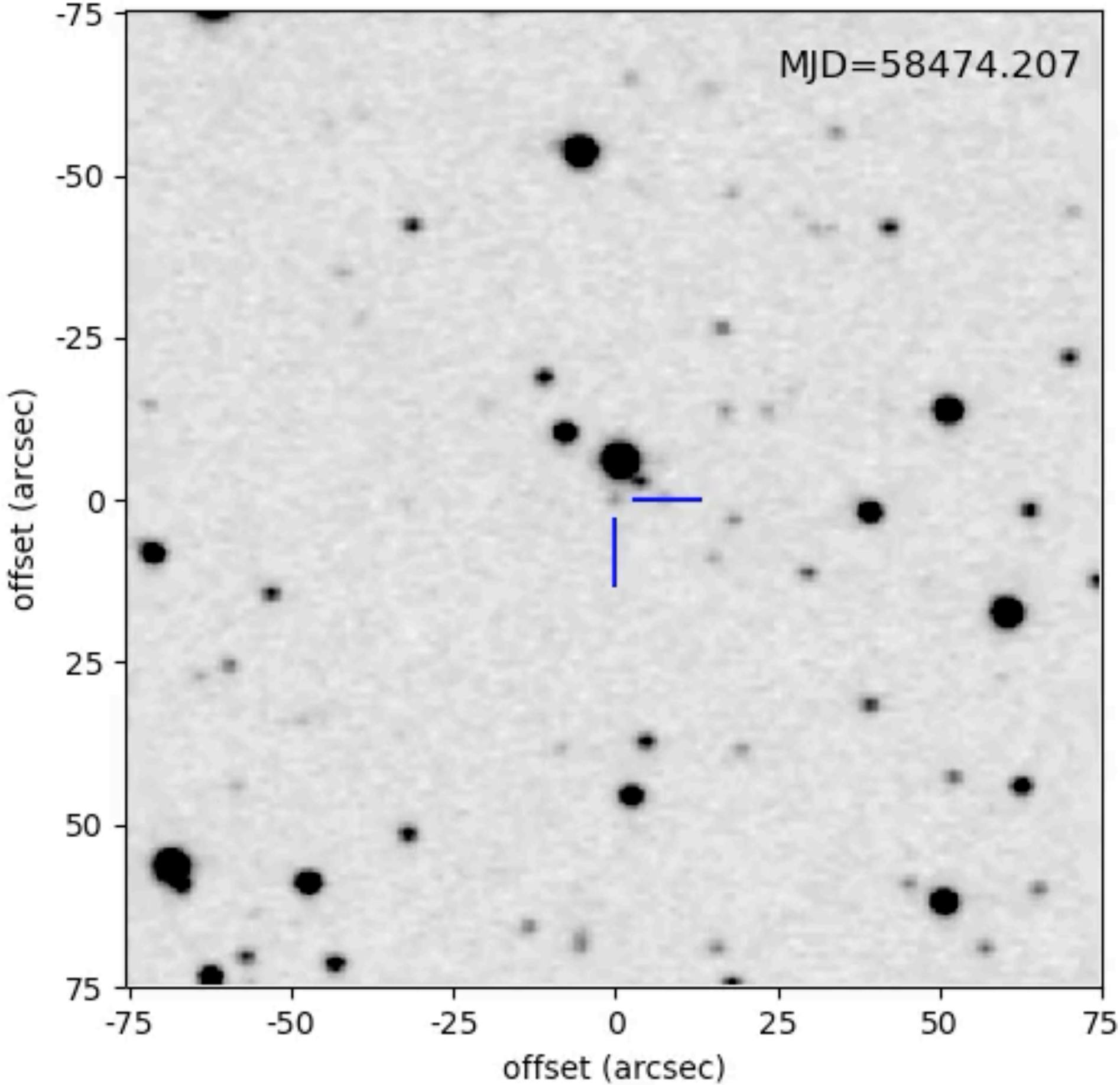
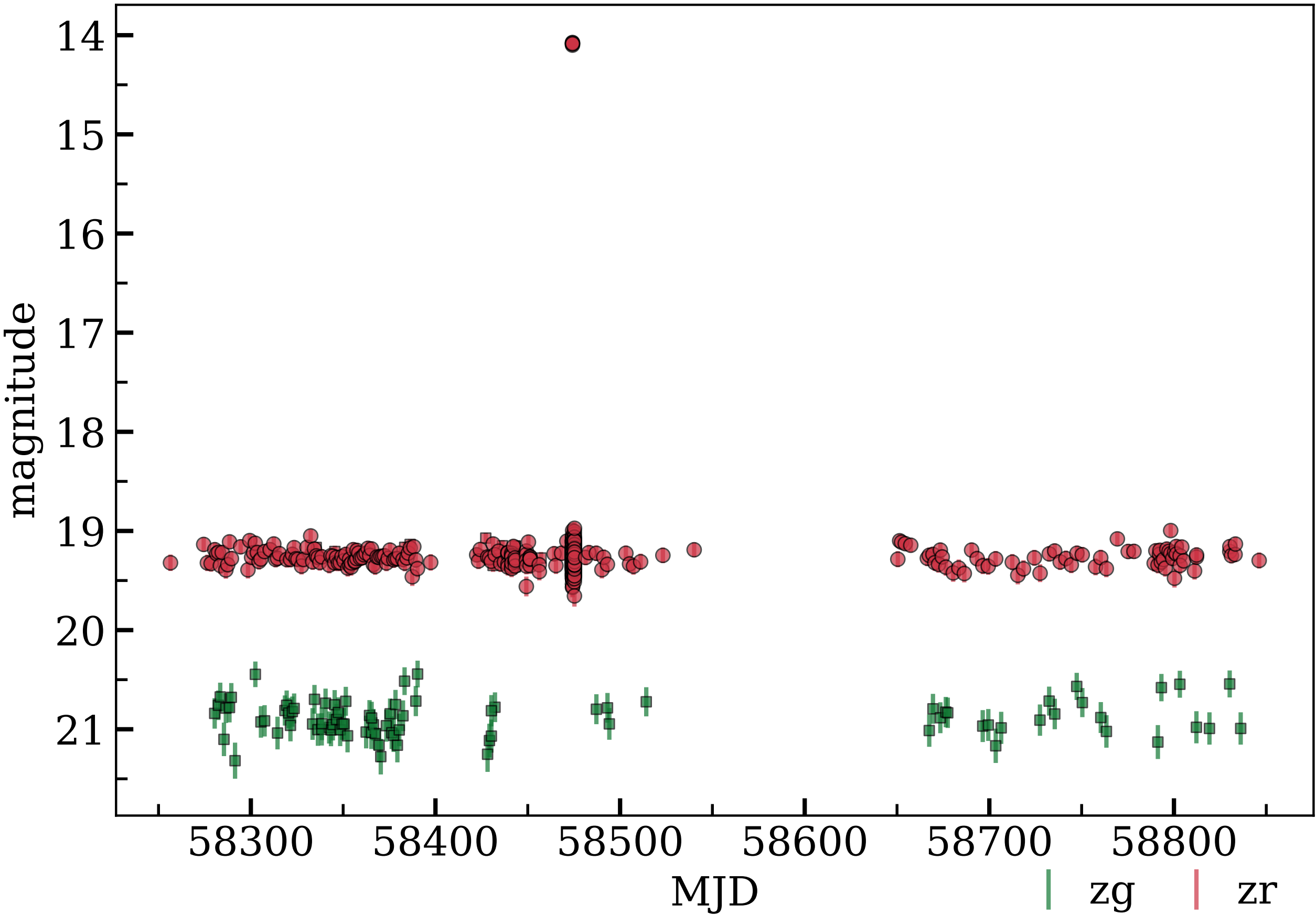


Sergey Lisakov

Barcelona

Asteroid 945

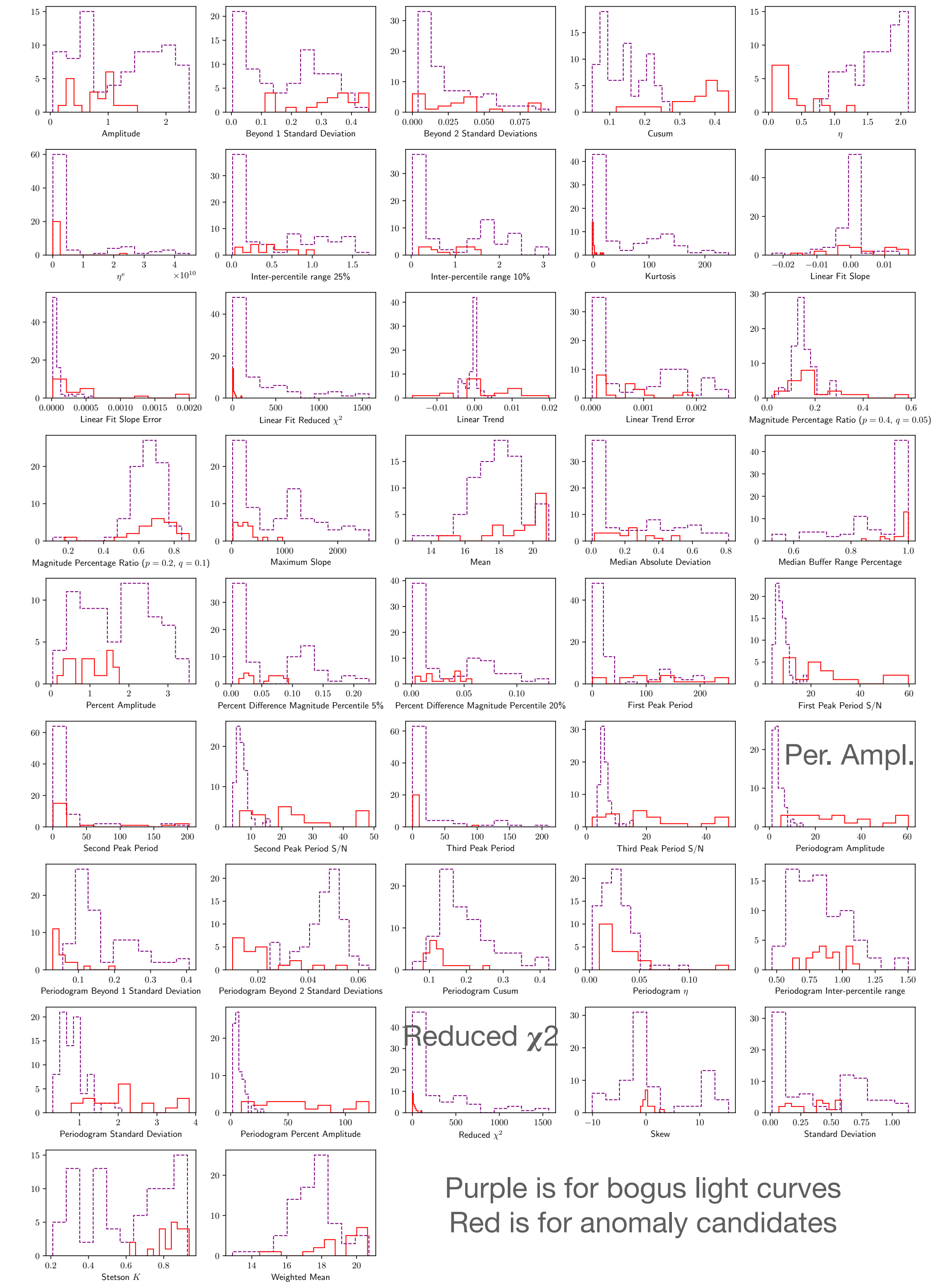
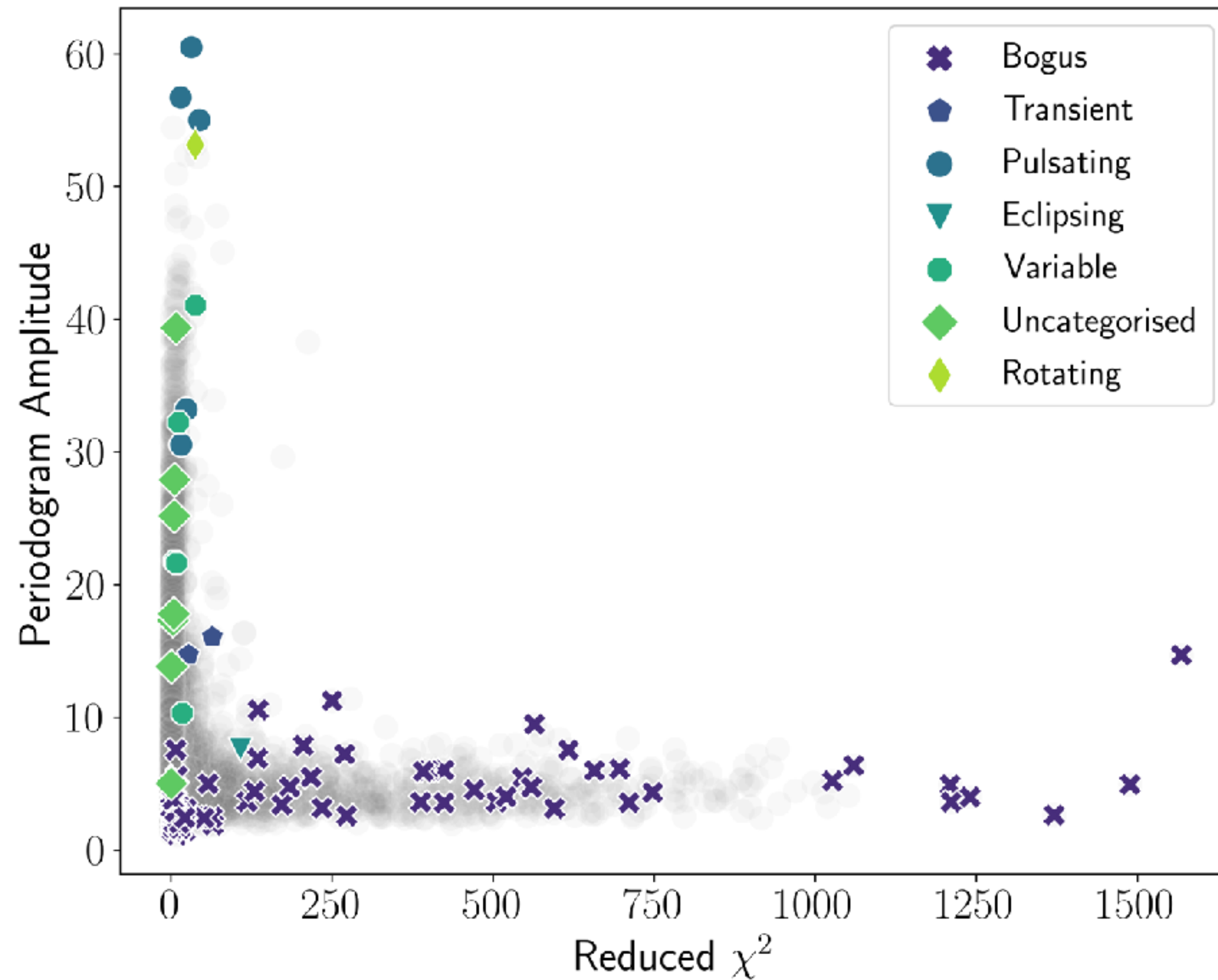
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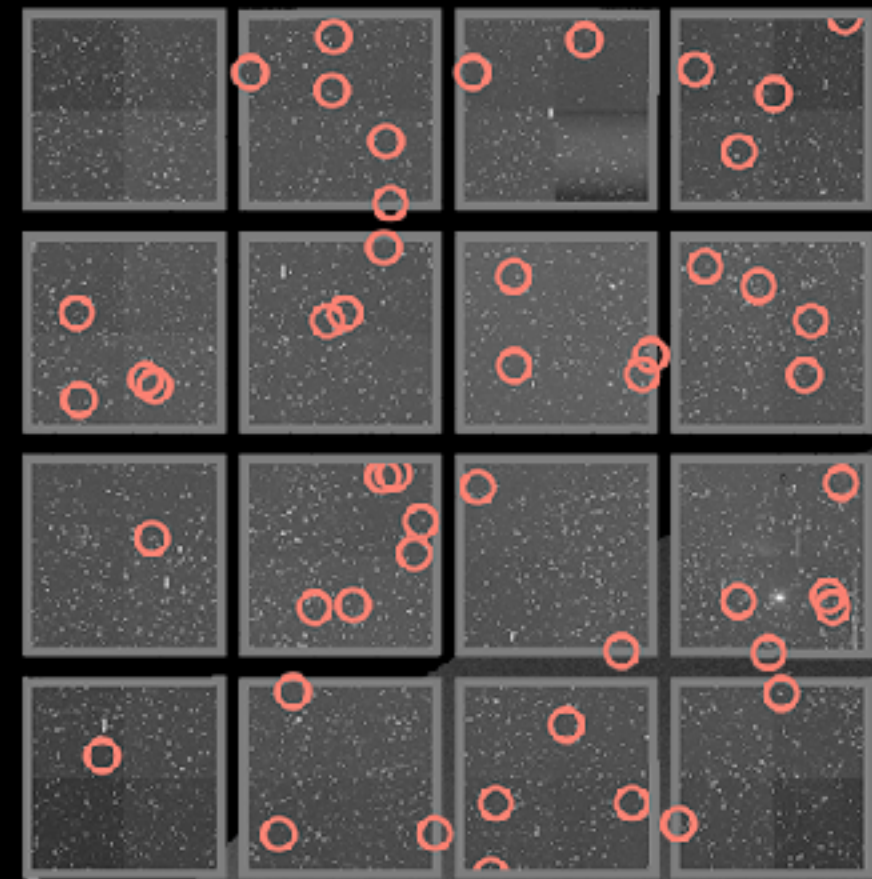


Exploratory Data Analysis

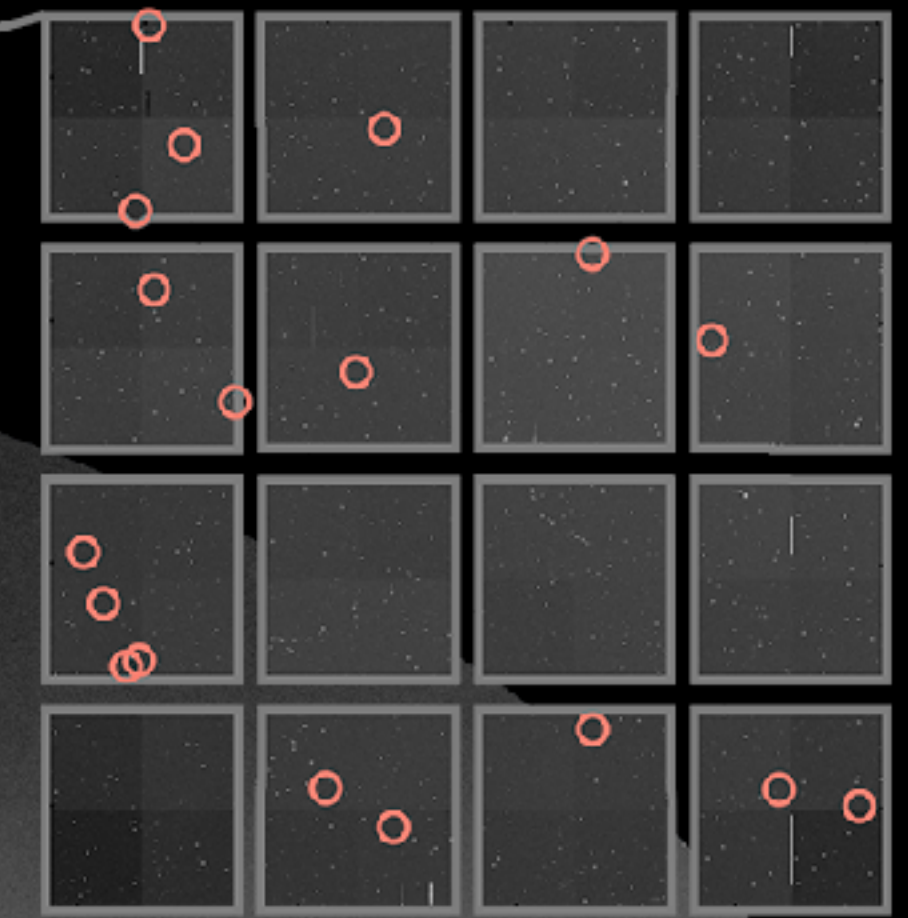
Bogus vs Real

M31 field



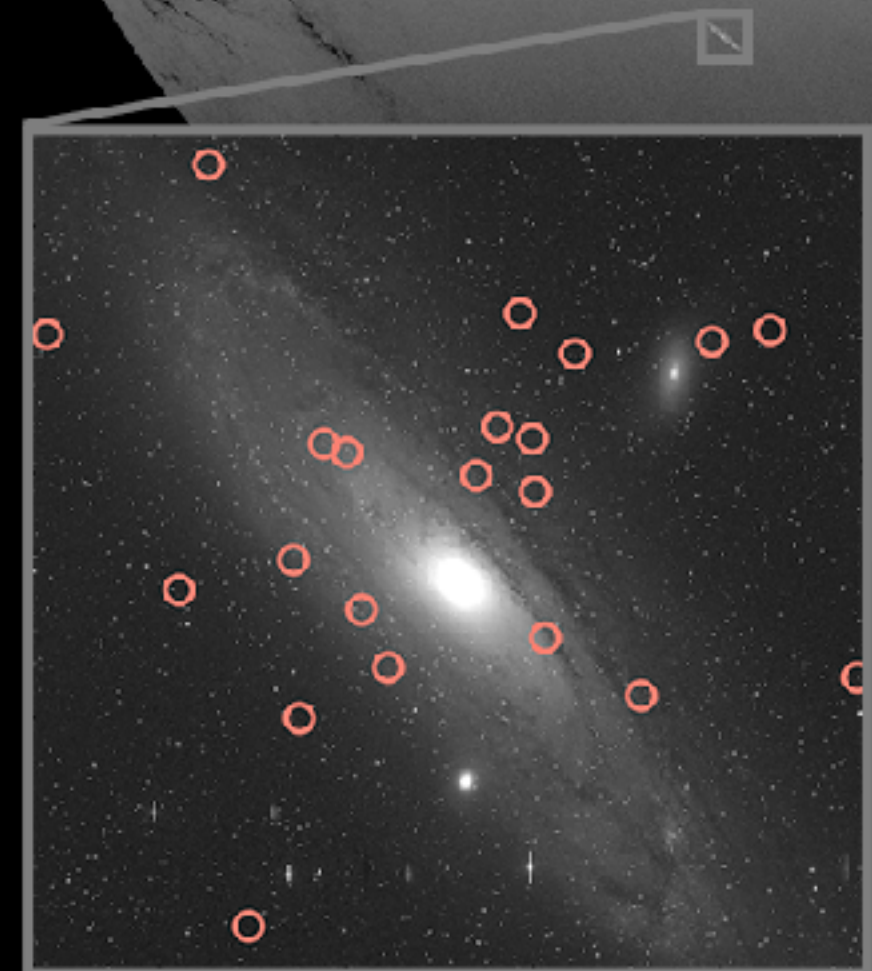


DISK



DEEP

Thank you!



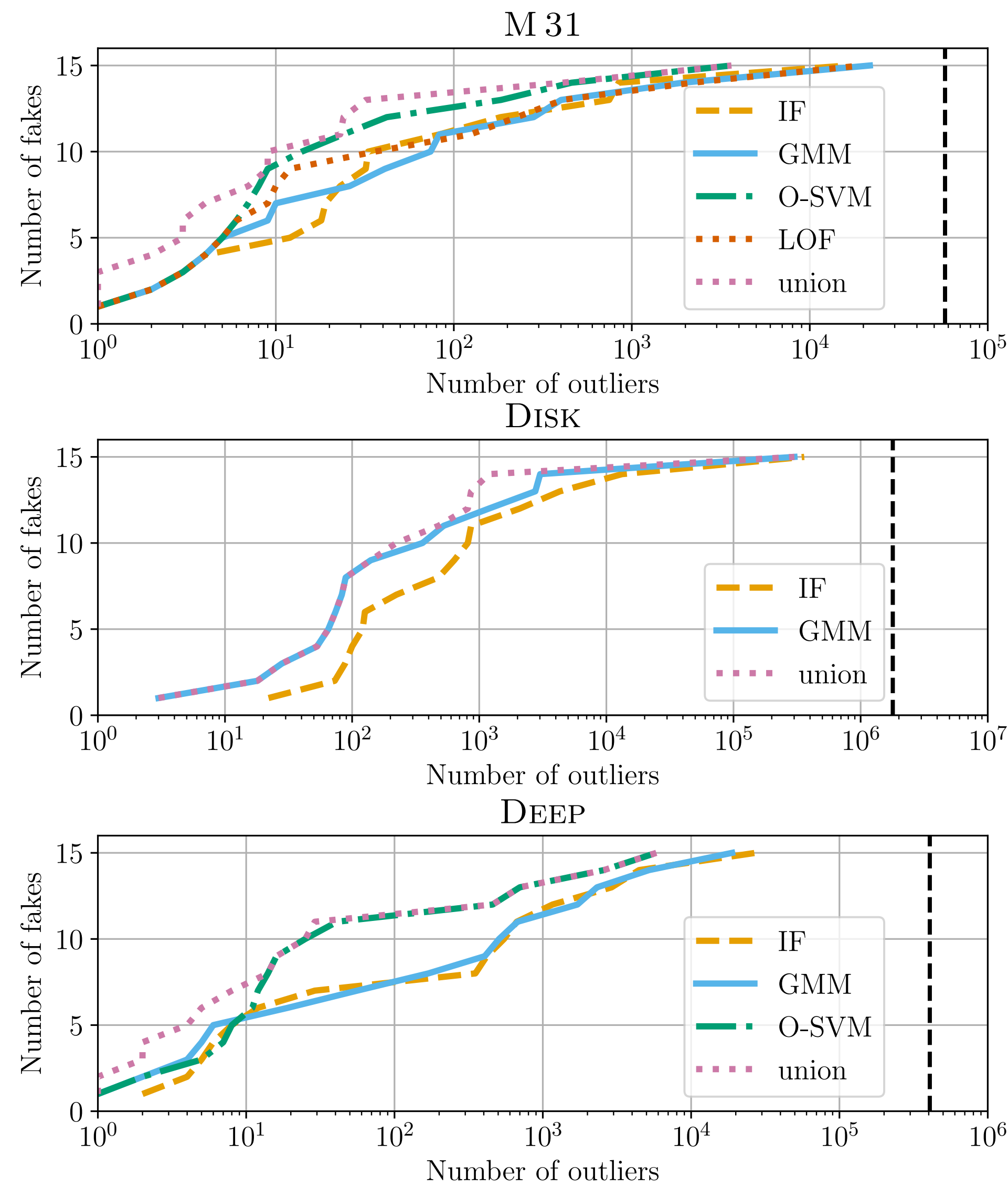
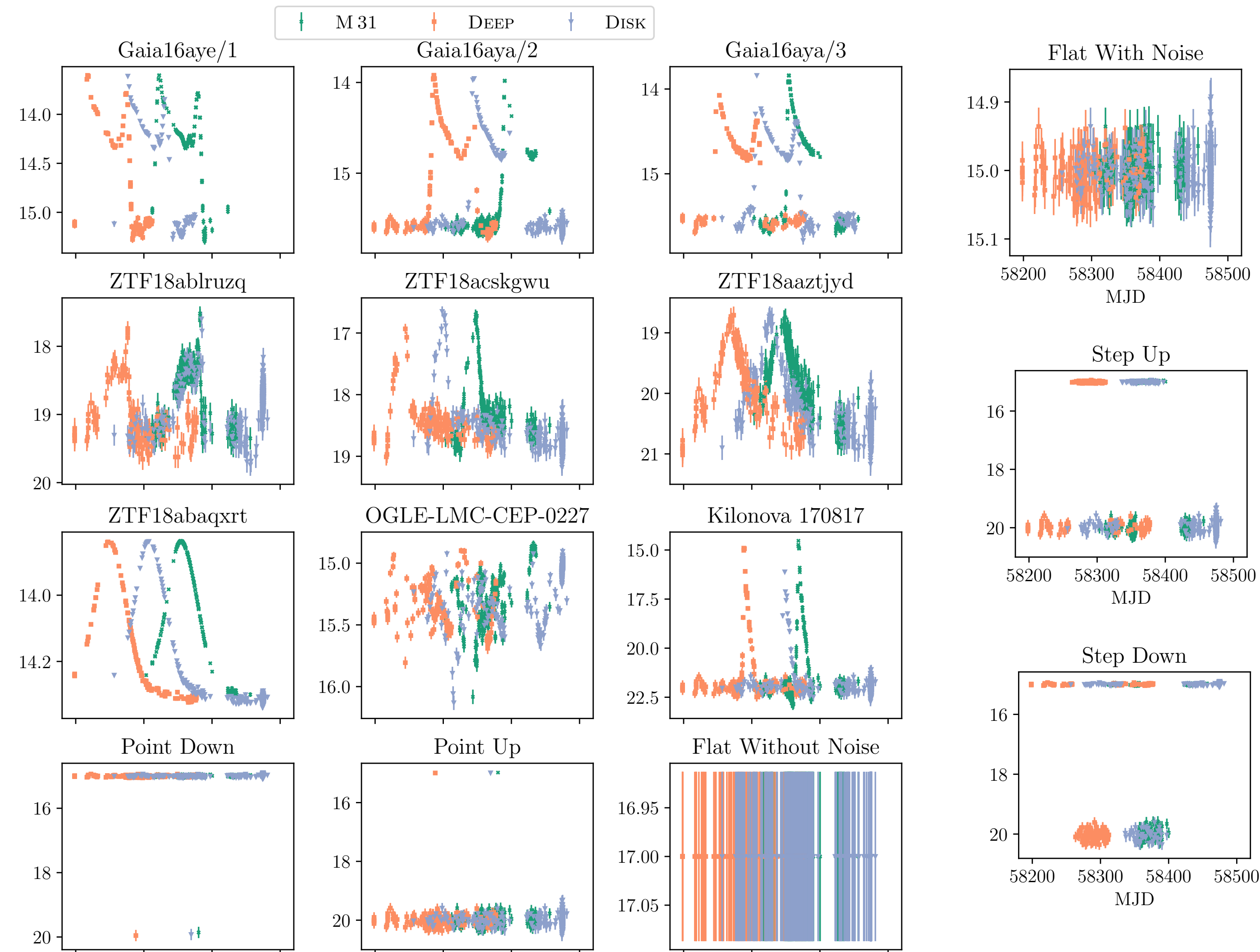
M31

 **ANOMALY
CANDIDATE**

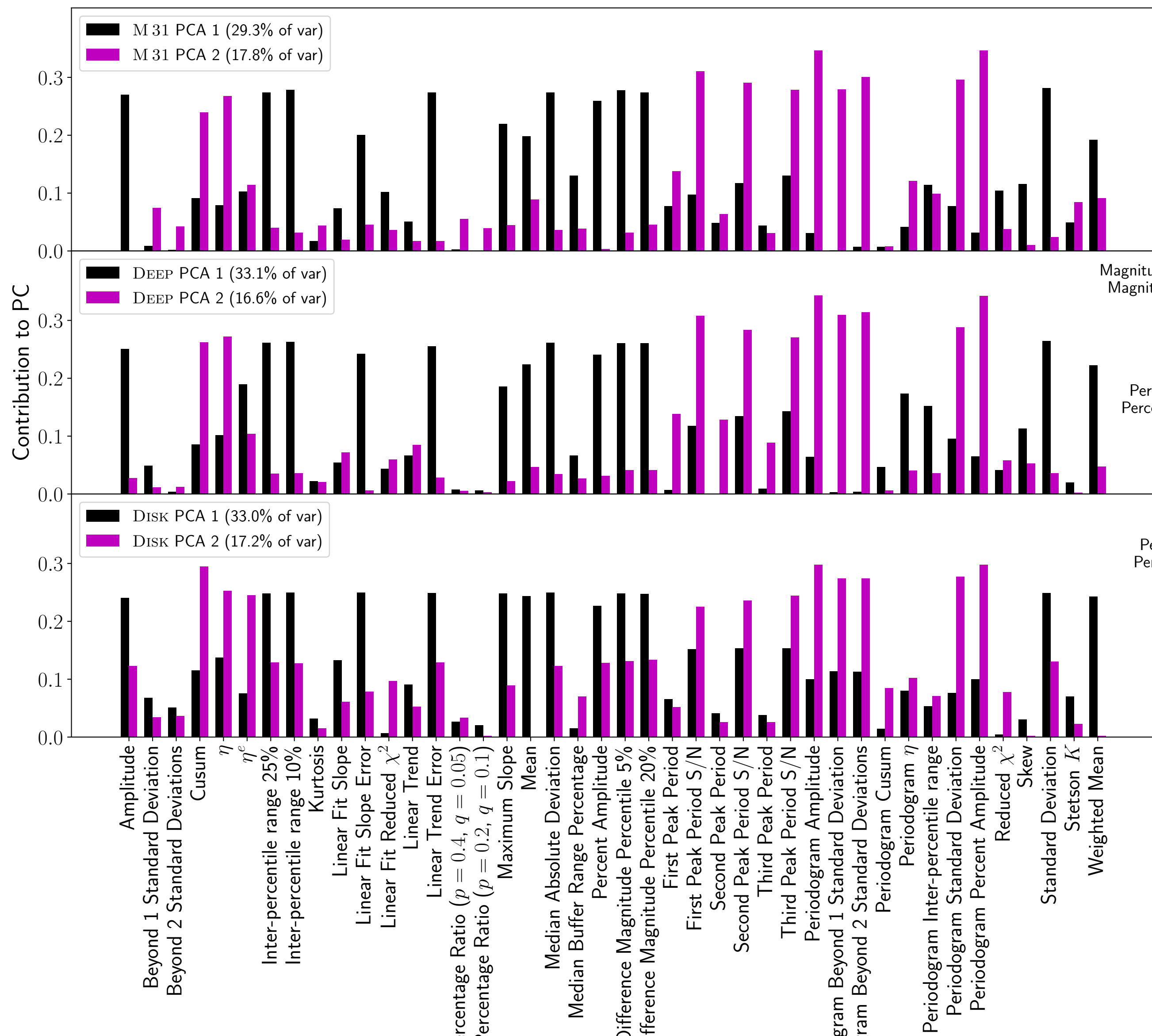
 **SNAD**

Outlier Detection Pipeline Verification

Fake light curves



Principal Component Analysis



Non-catalogued Anomaly Candidates

Three largest contributions to the three principal components

